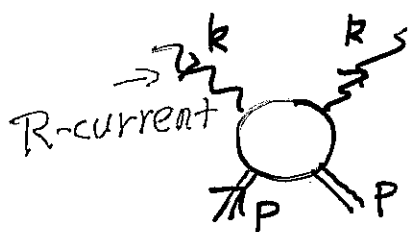


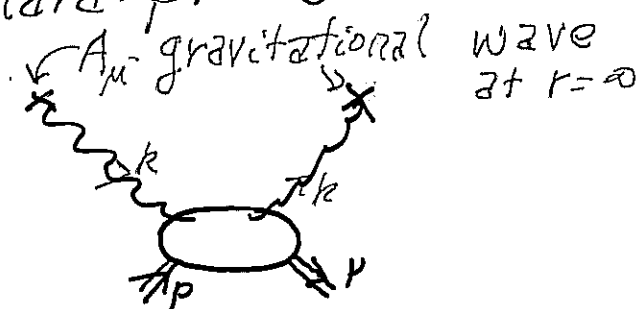
# 1. Review of Polchinski-Strassler

Study DIS on dilator, Break susy at  $r=r_0$ .

Use R-current as hard probe



Field Theory



Gravity

Gravitational wave induced at  $r=\infty$ , propagates in bulk and interacts with dilator which mostly lives at  $r=r_0$

On Field theory side use operator product expansion.

$$\chi_j = \frac{1}{2} - \sqrt{\frac{1}{2}} (j - j_0)$$

analogy of BFKL intercept  
 $\downarrow$   
 $j_0 = 2 - \frac{2}{\sqrt{11}}$

$$\chi_2 = 0 \quad \chi_j \underset{j=4,6,\dots}{\sim} -\frac{\lambda^{1/4}}{\sqrt{2}} (j - j_0)^{\frac{1}{2}} \quad \text{large and negative}$$

Strong violation of scaling! Partons

live only at small  $x$  at

$$x = \frac{1}{Q^2 r_0^2} e^{-\sqrt{11} y}$$

and below.

Difficult to get reasonable model of DIS for hadrons in AdS.

# DIS For a Plasma in AdS/CFT

with Y. Hatta  
E. Iancu

Object is to see what the  $N=4$  SYM plasma is made of at strong coupling.

A first step toward understanding hard probes of a strongly coupled plasma.

$N=4$  SYM  $\longleftrightarrow$  type IIB string theory on  $AdS_5 \times S^5$

$$g^2 = 4\pi g_s$$

$$\sqrt{\lambda} \equiv \sqrt{g^2 N_c} = R^2/\alpha'$$

$$ds^2 = \frac{r^2}{R^2} (-dt^2 + d\vec{x}^2) + \frac{R^2}{r^2} dr^2 + R^2 dS_5$$

At large  $g^2 N_c$  and small  $g^2$  string theory  $\rightarrow$  supergravity in  $AdS_5 \times S^5$

## 2. DIS on the strongly coupled $N=4$ SYM plasma

Goal is to evaluate

$$\text{Im} i \int d^4x e^{-ik \cdot x} \langle R(j_\mu(x) j_\nu(0)) \rangle_{\text{plasma}} = \frac{1}{T} W_{\mu\nu}(k)$$

at large  $k^2 = \vec{k}^2 - k_0^2$ , in order to see what the plasma is made of.

Do calculation on gravity side. Metric corresponding to plasma is

$$ds^2 = \frac{r^2}{R^2} [-f dt^2 + d\vec{x}^2] + \frac{R^2}{r^2} \frac{dr^2}{f}$$

with  $f = 1 - \frac{r_0^4}{r^4}$  and  $r_0 = \pi R^2 T$  the black hole horizon.

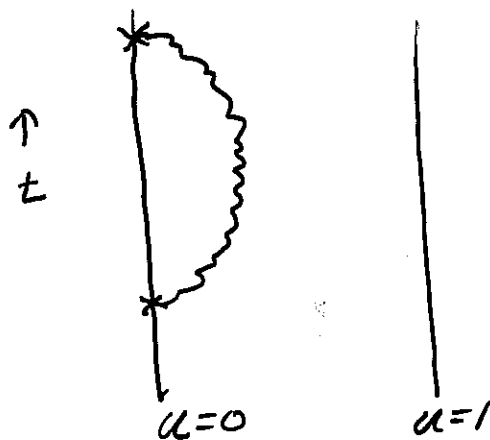
For technical reasons let  $u = r_0^2/r^2$ , then

$$ds^2 = \frac{(\pi R T)^2}{u} [-f dt^2 + d\vec{x}^2] + \frac{R^2}{4u^2 f} du^2 \quad f = 1 - u^2$$

Now  $r = \infty$  boundary is at  $u = 0$  and horizon is at  $u = 1$ .

Gravity Field induced  
by R-current obeys Yang-  
Mills equation

$$g^{\alpha\beta} \nabla_\alpha F_{\beta\gamma} = 0$$



Fix  $A_\mu = A_\mu^{(0)}$  on  $u=0$  boundary.  
Evaluate

$$S = \frac{-N_c^2}{64\pi^2 R} \int d^4x du \sqrt{-g} g^{\alpha\beta} g^{\gamma\delta} F_{\alpha\gamma} F_{\beta\delta}$$

and use, For  $A_\mu^{(0)}(x) = e^{ikz - i\omega t} A_\mu^{(0)}$ ,

$$R_{\mu\nu}(k) = -2 \frac{\delta}{\delta A_\mu^{(0)}} \frac{\delta}{\delta A_\nu^{(0)}} S \Big|_{A_\mu^{(0)}=0} \quad \frac{1}{T} W_{\mu\nu} = \text{Im} R_{\mu\nu}$$

In gauge where  $A_\mu = 0$ , one can write  
essential equation in terms of

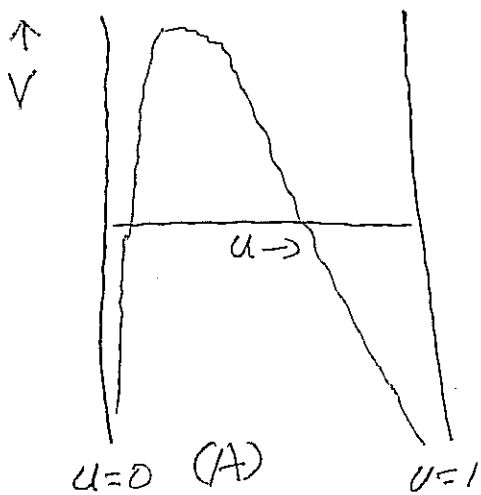
$$\psi(u) = \sqrt{u(1-u^2)} \frac{d}{du} A_0(u)$$

as

$$\psi''(u) = \frac{1}{u^2(1-u^2)^2} \left[ -\frac{1}{4} + Q^2 u - g^2 u^3 \right] \psi$$

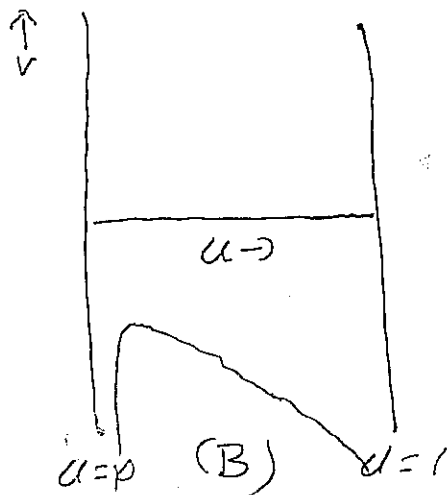
$$g = \frac{k}{2\pi T} \quad Q^2 = \left(\frac{1}{2\pi T}\right)^2 [k^2 - \omega^2]$$

$$V = \frac{1}{u^2(1-u^2)^2} \left[ -\frac{1}{4} + Q^2 u - g^2 u^3 \right]$$



$$\frac{g}{Q^3} \ll \frac{1}{4}$$

$$Q^2 \text{ large} \\ g/Q > 1$$



$$\frac{g}{Q^3} \gg \frac{1}{4}$$

(A) Large barrier so gravity wave lives near  $u=0$ . Imaginary part of correlator, coming from tunneling to  $u=1$ , is small.

(B) Gravity easily flows from  $u=0$  to  $u=1$  and is absorbed into black hole and hence a large  $W_{\mu\nu}$ .

$$S = -\frac{N^2}{32\pi^2} \left[ k A_c^{(0)} + \omega A_3^{(0)} \right]^2 \text{Volume} \cdot \text{time} \left[ \frac{2}{3}(\gamma + \ln \frac{g}{3}) - i\frac{\pi}{3} \right]$$

$$\Rightarrow \langle J_\mu J_\nu \rangle_k \sim \left( \frac{k_\mu k_\nu}{k^2} + g_{\mu\nu} \right) k^2 N^2$$

$$\left[ \frac{1}{x} \frac{dF_2}{d^3b} \sim k^2 T N^2 \right]$$

### 3. Interpretations

#### Partonic

Coherence time for  $\langle JJ \rangle$  correlator is

$$t_c \simeq \frac{2k}{K^2} = \frac{1}{X\pi^2} \quad X = \frac{K^2}{2KT}$$

length of matter involved is  $\Delta z = t_c = \frac{1}{X\pi^2}$

$$\frac{dF_2^{\Delta z}}{d^2b} = X \cdot \underbrace{\frac{1}{X} \frac{dF_2}{d^3b}}_{K^2 N^2} \cdot \underbrace{\frac{1}{X\pi^2}}_{\Delta z} = K^2 N^2 \quad \frac{K}{K^3} > \frac{1}{T} \text{ or } \underline{\underline{\frac{1}{X} > \frac{K}{T}}}$$

To see partons boost material by  $\eta$   
where  $\cosh \eta = (K/T)^{1/X}$

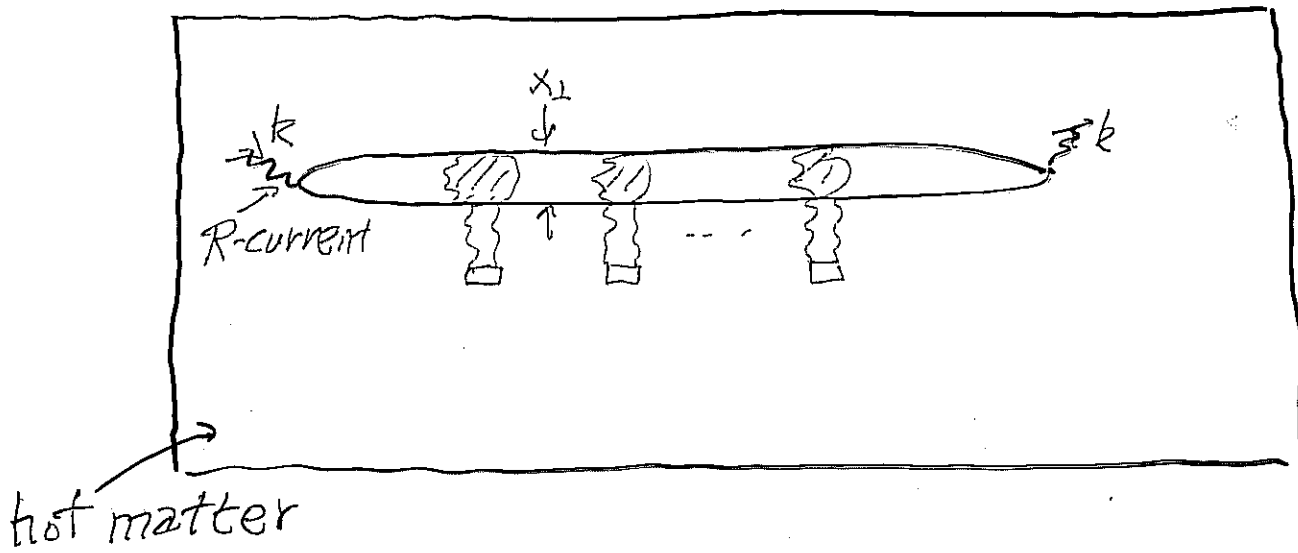
Then  $\Delta z' \simeq 1/K$  and current measures  
one layer of partons in  $\Delta z'$

$$\frac{dF_2^{\Delta z'}}{d^2z} \sim N^2 K^2 \quad (\text{saturation})$$

$X < TK$

When measured at scale  $K$  all  
energy is located in partons with  $X = TK$ .

# Multiple scattering



In multiple scattering usual Factor

$$\text{rate} = \int \sigma \sim \int x_{\perp}^2 X G(x, 1/x_{\perp}^2) \frac{\alpha N_c}{N_c^2 - 1}$$

$$\downarrow$$

$$\int x_{\perp}^2 \frac{c}{x} \quad \frac{1}{x} = kT x_{\perp}^2$$

Now real, not absorptive potential.  $1/x$  Factor determined by operator product expansion with  $\mathcal{O}_{pp}$  as only allowed operator.

Potential Forces  $x_{\perp}$  to grow with passage through material? When  $x_{\perp} \sim 1/T$  interaction becomes strong.