

# **Indication for deconfinement at RHIC**

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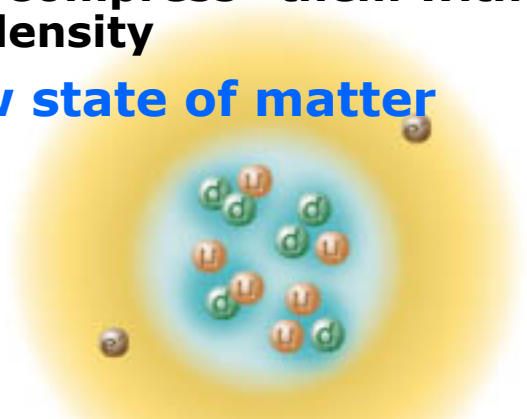
## **Buda-Lund hydro fits to spectra and HBT radii**

**M. Csanád, T. Csörgő, B. Lörstad and A. Ster**  
**(Budapest & Lund)**

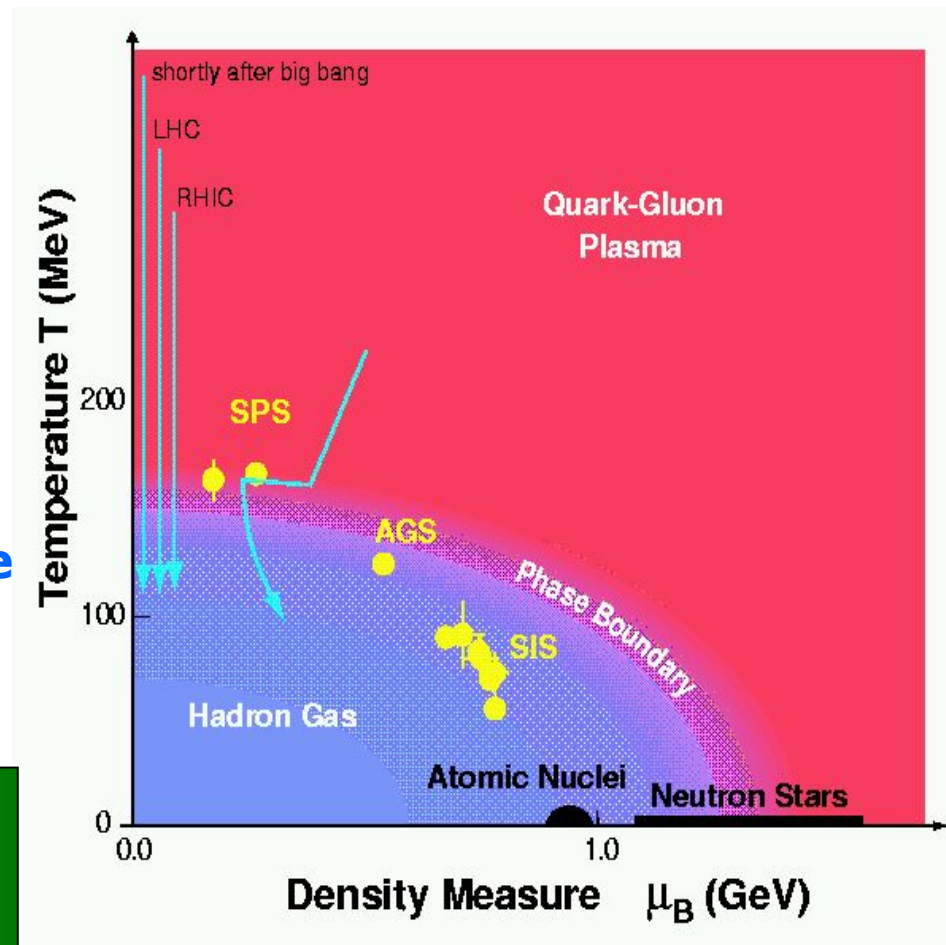
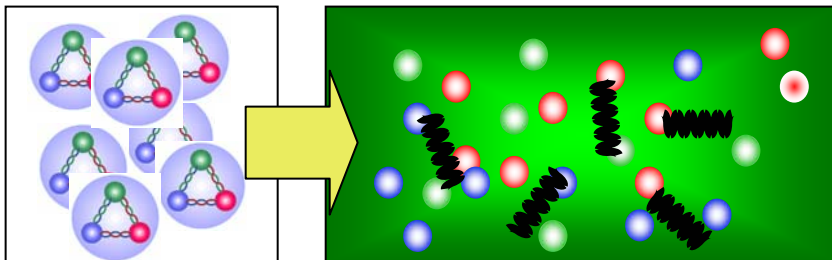
- Buda-Lund hydro model
- Buda-Lund fits to final RHIC Au+Au data
- A resolution of the “RHIC HBT puzzle”
  - Indication for a hot center,  $T > T_c$
  - Confirmation by independent measurement
  - Prediction for new observables

# Phases of QCD Matter

- **Quark Gluon Plasma**
  - “Ionize” nucleons with heat
  - “Compress” them with density
- **New state of matter**

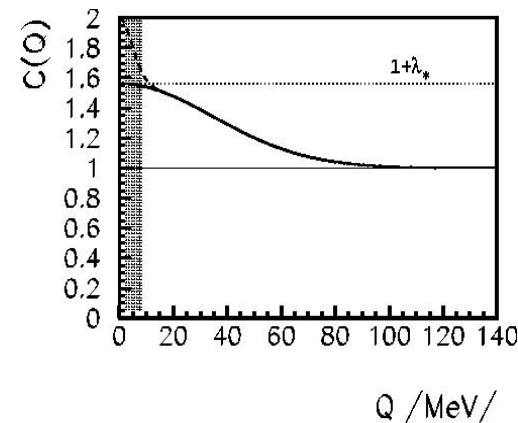
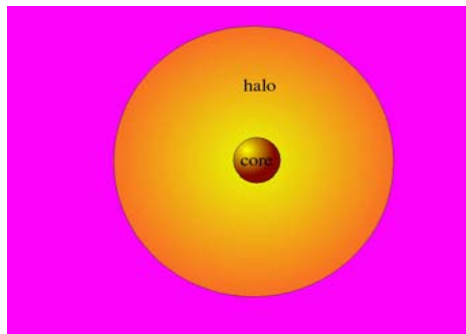


**Z. Fodor and S.D. Katz:**  
 $T_c = 172 \pm 3$  MeV even at finite  
baryon density,  
Cross over like transition.  
(hep-lat/0106002)



# Principles for Buda-Lund hydro model

- Analytic expressions for all the observables
- 3d expansion, local thermal equilibrium, symmetry
- Goes back to known hydro solutions in nonrel limit
- Separation of the Core and the Halo
  - Core: hydrodynamic evolution
  - Halo: decay products of long-lived resonances



# Buda-Lund hydro model

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**The general form of the emission function:**

$$S_c(x, p)d^4x = \frac{g}{(2\pi)^3} \frac{p^\mu d^4\Sigma_\mu(x)}{\exp\left(\frac{p^\nu u_\nu(x)}{T(x)} - \frac{\mu(x)}{T(x)}\right)} + s_q$$

**Calculation of observables with core-halo correction:**

$$N_1(p) = \frac{1}{\sqrt{\lambda_*}} \int d^4x S_c(p, x)$$
$$C(Q, p) = 1 + \left| \frac{\tilde{S}(Q, p)}{\tilde{S}(0, p)} \right|^2 = 1 + \lambda_* \left| \frac{\tilde{S}_c(Q, p)}{\tilde{S}_c(0, p)} \right|^2$$

**Assuming special shapes for the flux, temperature, chemical potential and flow:**

# Buda-Lund hydro model

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**Invariant single particle spectrum:**

$$N_1 = \frac{d^2n}{2\pi m_t dm_t dy} = \frac{g}{(2\pi)^3} \overline{E} \overline{V} \overline{C} \frac{1}{\exp\left(\frac{p^\mu u_\mu(x_s) - \mu(x_s)}{T(x_s)}\right) + s_q}$$

**Invariant Buda-Lund correlation function: oscillating, non-Gaussian prefactor!**

$$C_2(k_1, k_2) = 1 + \lambda_* \Omega(Q_{||}) \exp\left(-Q_{||}^2 R_{||}^2 - Q_{\pm}^2 R_{\pm}^2 - Q_{\perp}^2 R_{\perp}^2\right)$$

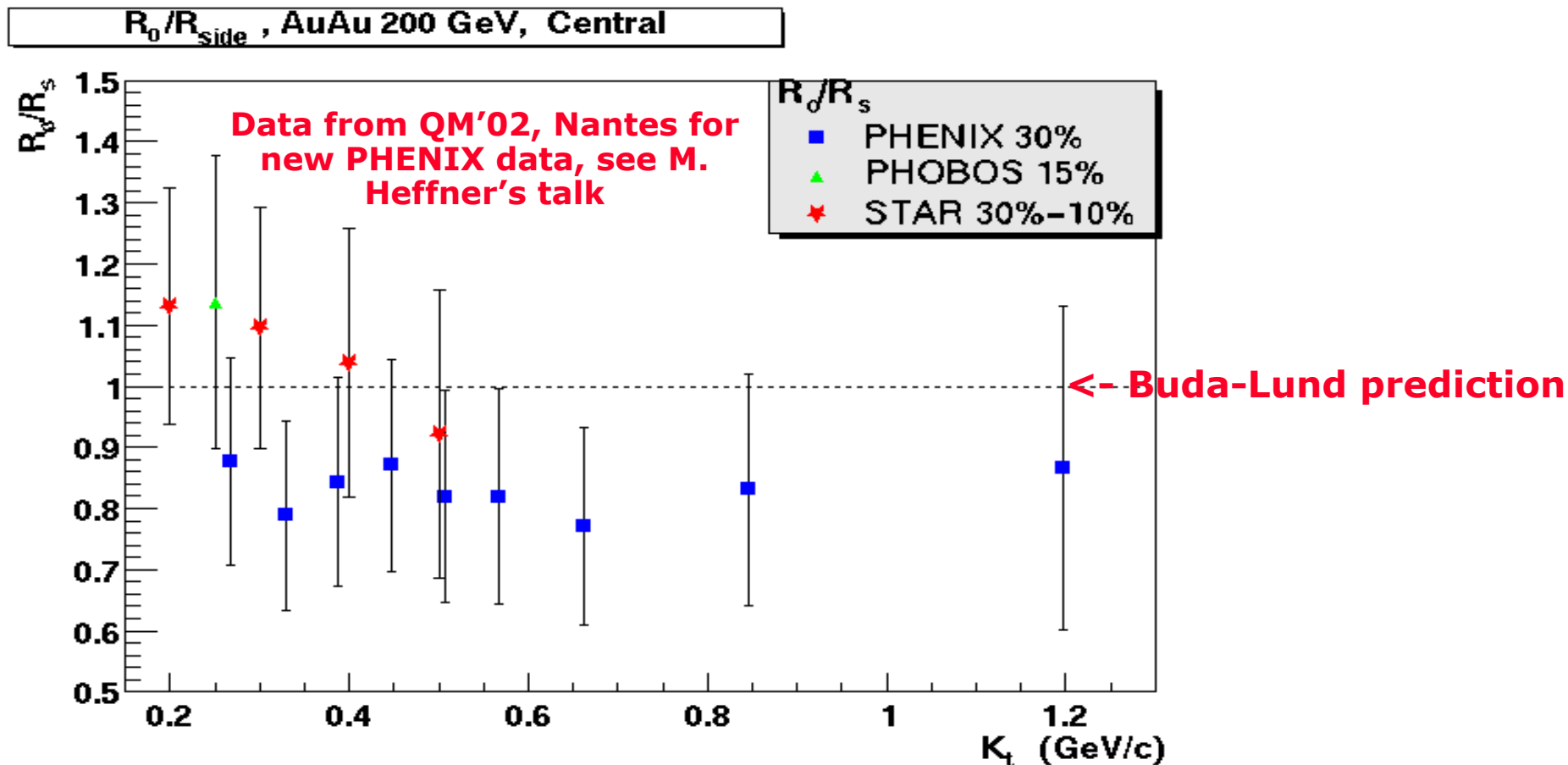
**Non-invariant Bertsch-Pratt parameterization, Gaussian approximation:**

$$C_2(k_1, k_2) = 1 + \lambda_* \exp\left(-Q_o^2 R_o^2 - Q_s^2 R_s^2 - Q_l^2 R_l^2 - 2Q_{os}^2 R_o R_s\right)$$

**Non-Gaussian BL form  $\longrightarrow$  Gaussian BP *approximation*:**

$$R_{||,\Omega}^2 = R_{||}^2 \left(1 + \frac{\overline{\Delta\eta}^2}{\overline{\eta}}\right)$$

# $R_{\text{out}}/R_{\text{side}}$ ratios at 200 GeV



- This reflects **symmetry** of the flow!  
(T. Csörgő & B. Lörstad, PRC54(1996)1390, NPA590(1995)465)
- Note also:  $R_{\text{side}}/R_{\text{long}} \rightarrow \text{const} (\sim 1)$  as  $m_t \gg T_0$  is another BL prediction
- **Predicted** by a dynamical model with supercooling, hep-ph/940636

# Buda-Lund fits to RHIC data

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## Calculation program packages:

**Version 1.0:** Calculation of IMD & Radii, linearized saddle-point equations  
(NA22, NA44, NA49, prel. RHIC run-1)

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**Version 1.4:** Calculation for  $dN/\Delta\eta$  added

**Version 1.5:** Saddle point found exactly in direction  $\eta$   
<http://www.kfki.hu/~csorgo/budalund/>

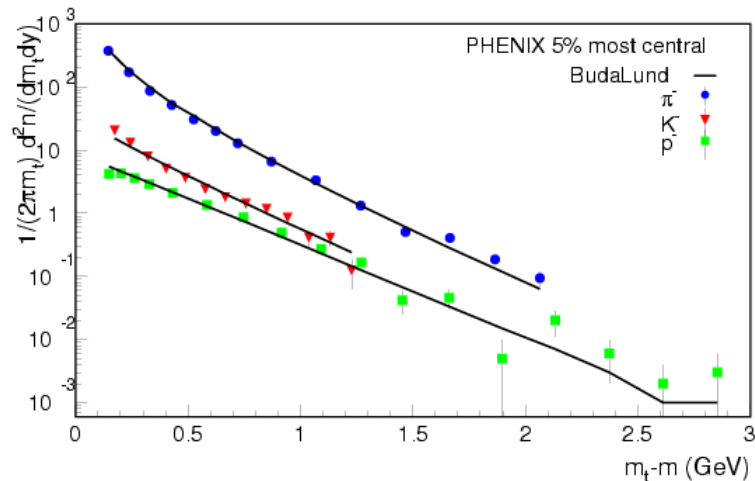
**Version 2.0:** Exact analytic calculation (in progress)

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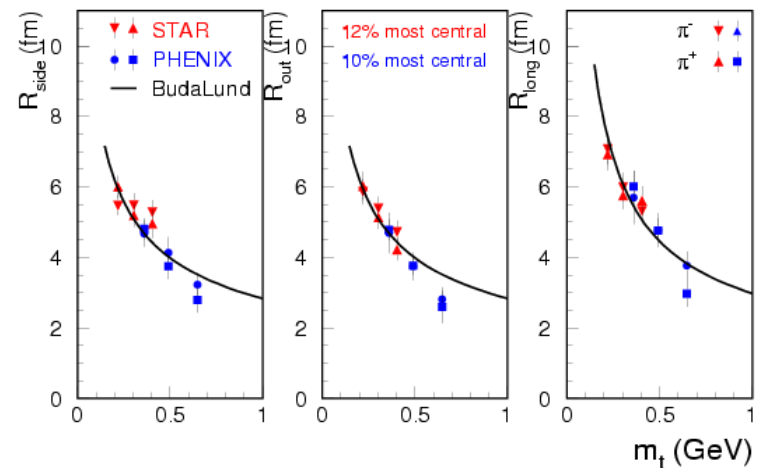
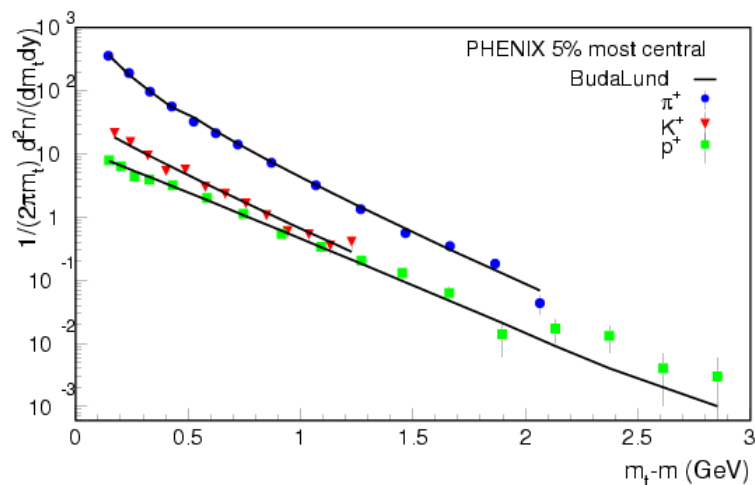
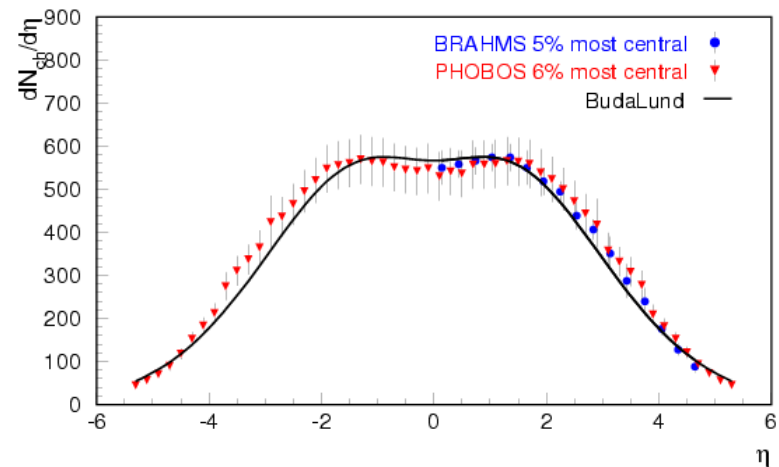
**All versions:** Analytic expressions for all observables  
and **simultaneous fits** to all the observables

# BudaLund fits, 130 GeV RHIC data

BudaLund hydro fits to 130 AGeV Au+Au



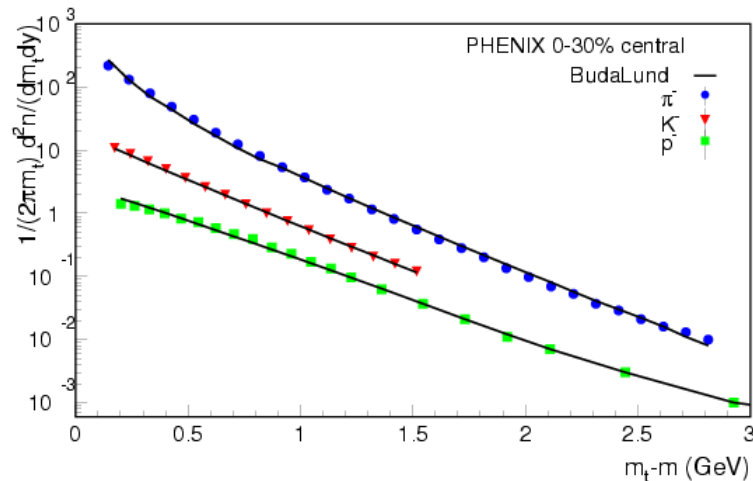
BudaLund hydro fits to 130 AGeV Au+Au



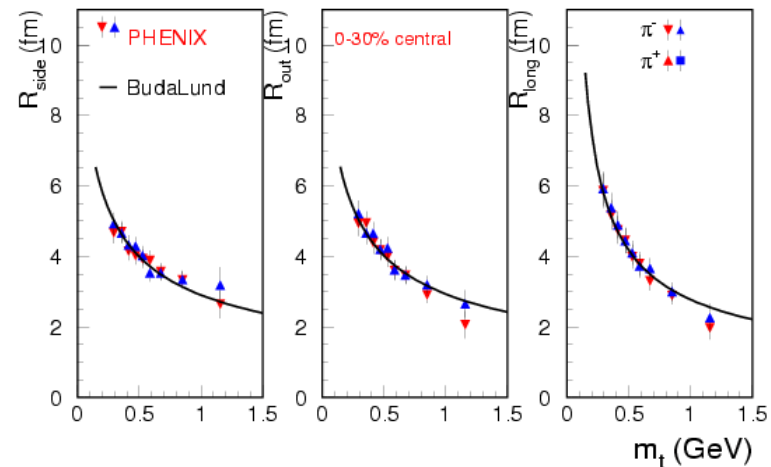
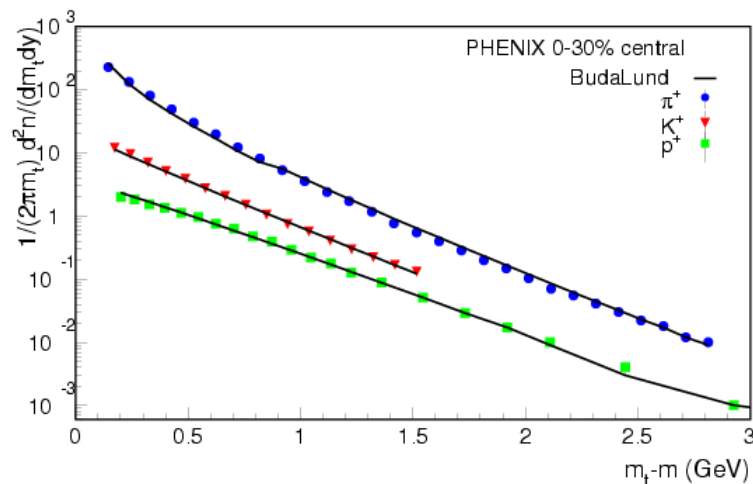
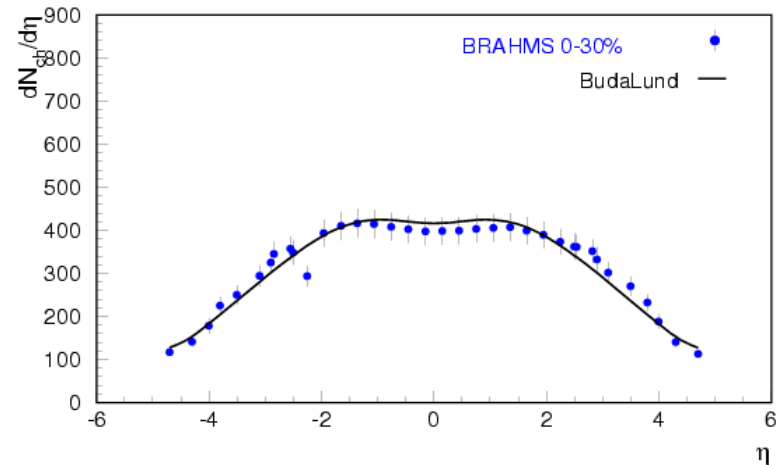


# BudaLund fits, 200 GeV RHIC data

BudaLund v1.5 hydro fits to 200 AGeV Au+Au



BudaLund v1.5 fits to 200 AGeV Au+Au



# Comparing RHIC Au+Au to SPS results

BL v1.5 parameters	RHIC 200 GeV Au+Au	RHIC 130 GeV Au+Au	Pb+Pb SPS	h+p SPS
$T_0$ [MeV]	<b>200 <math>\pm</math> 9</b>	<b>214 <math>\pm</math> 7</b>	<b>139 <math>\pm</math> 6</b>	<b>140 <math>\pm</math> 3</b>
$\langle u_t \rangle$	<b>1.5 <math>\pm</math> 0.1</b>	<b>1.0 <math>\pm</math> 0.1</b>	<b>0.55 <math>\pm</math> 0.06</b>	<b>0.20 <math>\pm</math> 0.07</b>
$R_s$ [fm]	<b>11.6 <math>\pm</math> 1</b>	<b>8.6 <math>\pm</math> 0.4</b>	<b>7.1 <math>\pm</math> 0.2</b>	<b>0.88 <math>\pm</math> 0.13</b>
$T_{\text{surf}}$ [MeV]	<b>0.5 <math>T_0</math> fixed</b>	<b>0.5 <math>T_0</math> fixed</b>	<b>131 <math>\pm</math> 8</b>	<b>82 <math>\pm</math> 7</b>
$\tau_0$ [fm/c]	<b>5.7 <math>\pm</math> 0.2</b>	<b>6.0 <math>\pm</math> 0.2</b>	<b>5.9 <math>\pm</math> 0.6</b>	<b>1.4 <math>\pm</math> 0.1</b>
$\Delta\tau$ [fm/c]	<b>1.9 <math>\pm</math> 0.5</b>	<b>0.3 <math>\pm</math> 1.2</b>	<b>1.6 <math>\pm</math> 1.5</b>	<b>1.3 <math>\pm</math> 0.3</b>
$\Delta\eta$	<b>3.1 <math>\pm</math> 0.05</b>	<b>2.3 <math>\pm</math> 0.4</b>	<b>2.1 <math>\pm</math> 0.4</b>	<b>1.36 <math>\pm</math> 0.02</b>
$T_{\text{evap}}$ [MeV]	<b>127 <math>\pm</math> 13</b>	<b>102 <math>\pm</math> 11</b>	<b>87 <math>\pm</math> 24</b>	-
$\mu_0^\pi$ [MeV]	<b>-2 <math>\pm</math> 14</b>	<b>63 <math>\pm</math> 11</b>		
$\mu_0^{K^+}$ [MeV]	<b>16 <math>\pm</math> 19</b>	<b>98 <math>\pm</math> 19</b>		
$\mu_0^{p^-}$ [MeV]	<b>97 <math>\pm</math> 28</b>	<b>315 <math>\pm</math> 27</b>		
$\mu_B$ [MeV]	<b>61 <math>\pm</math> 39</b>	<b>77 <math>\pm</math> 38</b>		
$\chi^2/\text{NDF}$	<b>126/208=0.61</b>	<b>158/180=0.9</b>		
CL	<b>100 %</b>	<b>88 %</b>		

A  $5\sigma$  effect,  $T_0 > T_c$

$T_0$  (RHIC)  $>$   $T_0$  (SPS)

Indication for quarks & hard EOS

$\mu_B = 77 \pm 38$  MeV

# Comparison of expansion rates:

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## RHIC and the universe

Universality of the Hubble expansion:  $u = H r$

Hubble constant of the Universe:  $H_0 = (71 \pm 7) \text{ km/sec/Mpc}$

converted to SI units:  $H_0 = (2.3 \pm 0.2) \times 10^{-18} \text{ sec}^{-1}$

Hubble constant at Au+Au collisions with 200 GeV

Method a)  $H_{\text{RHIC}} = \langle u_t \rangle / R_G \sim (3.8 \pm 0.5) \times 10^{22} \text{ sec}^{-1}$

Method b)  $H_{\text{RHIC}} = 1/\tau_0 \sim (5.1 \pm 0.1 \times 10^{22}) \text{ sec}^{-1}$

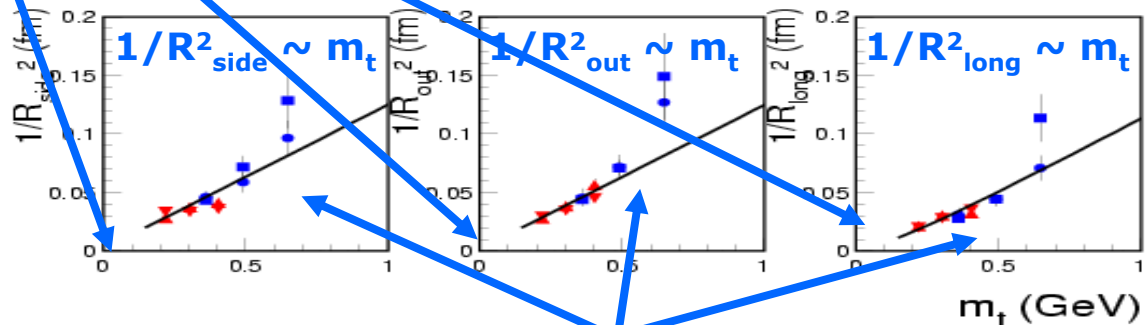
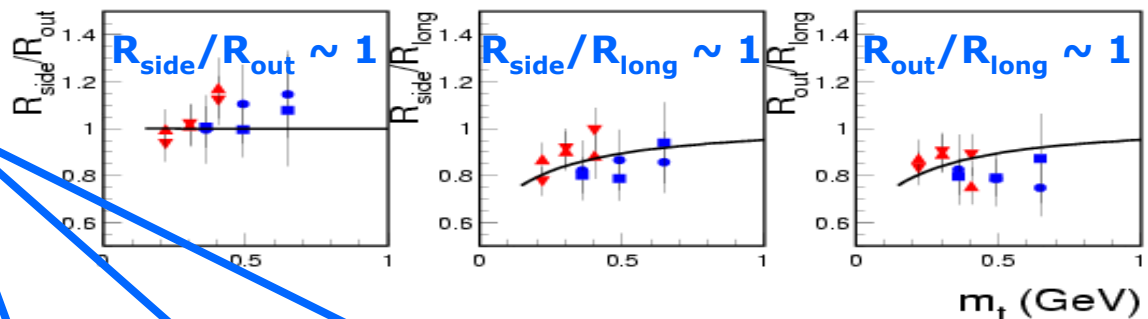
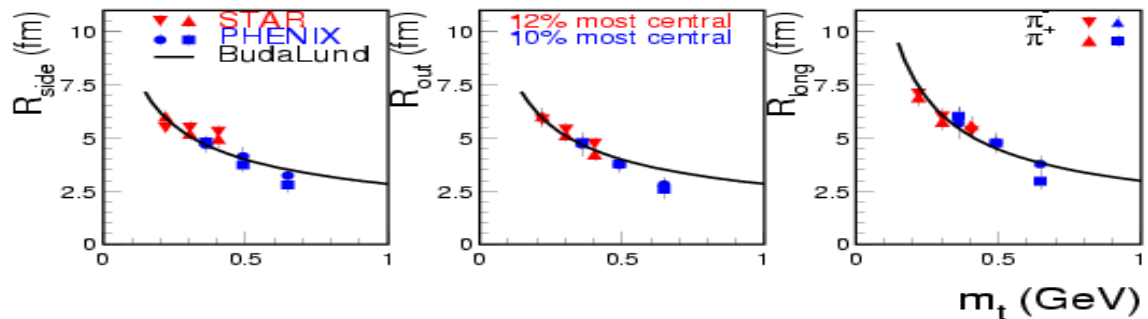
Ratio of expansion rates:

$$H_{\text{RHIC}} / H_0 \sim 2 \times 10^{40}$$

approx. the ratio of the ages of the objects,  
without correction for inflation...

# New ways of plotting the results

BudaLund hydro fits to 130 AGeV Au+Au



same slopes  $\sim$  fully developed, 3d Hubble flow

$$1/R_{\text{eff}}^2 = 1/R_{\text{geom}}^2 + 1/R_{\text{therm}}^2$$

and  $1/R_{\text{therm}}^2 \sim m_t$



intercept is nearly 0,  
indicating  $1/R_G^2 \sim 0$ ,

thus  $\mu(x)/T(x) = \text{const!}$

reason for success of  
thermal models @ RHIC!

# SUMMARY

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The analysis of RHIC Au+Au data from BRAHMS, PHOBOS, PHENIX and STAR show:

- **successful** Buda-Lund hydro fits

(also at h+p and Pb+Pb at SPS)

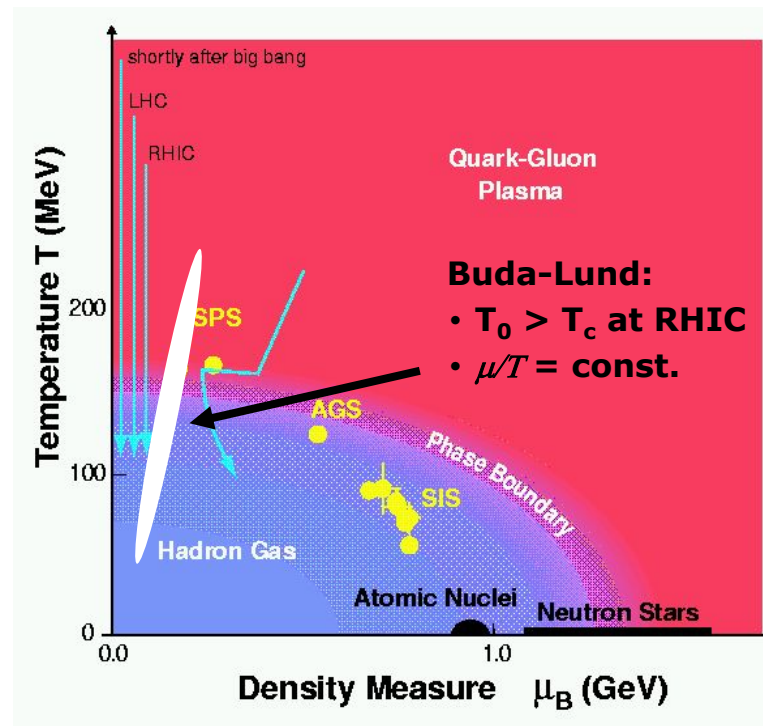
- **indication** for deconfinement at RHIC

( $T > T_c = 172 \text{ MeV}$  by  $3\sigma$  at RHIC, but not at SPS)

- **evidence** for a developed 3d Hubble flow at RHIC

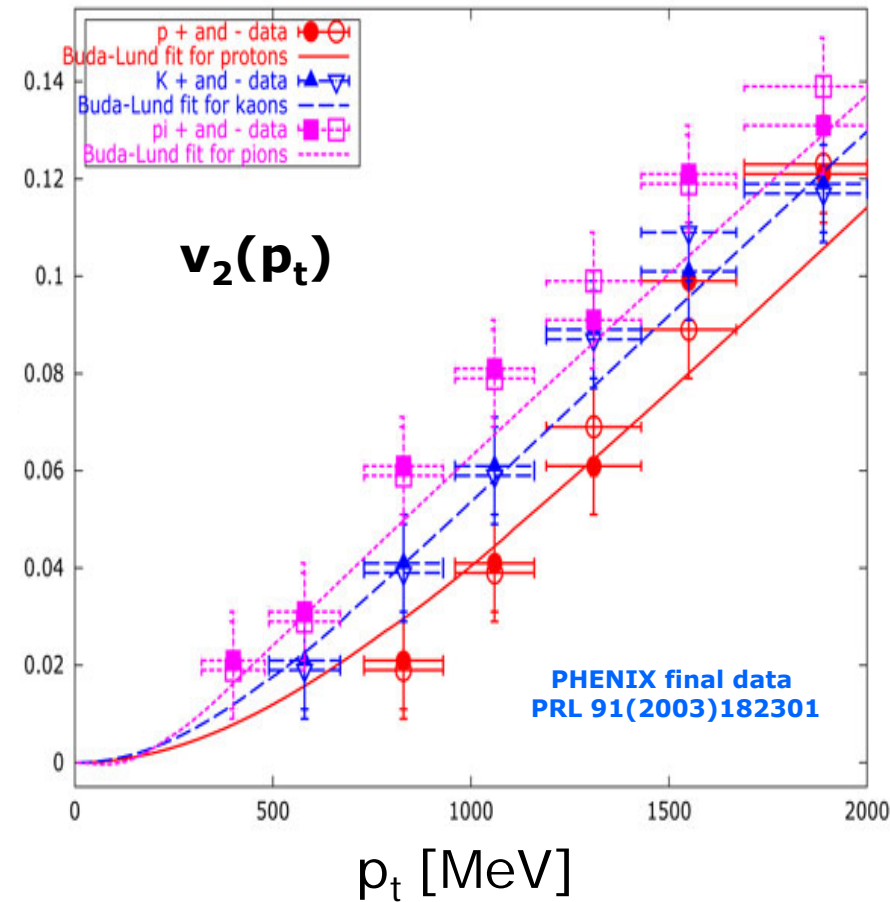
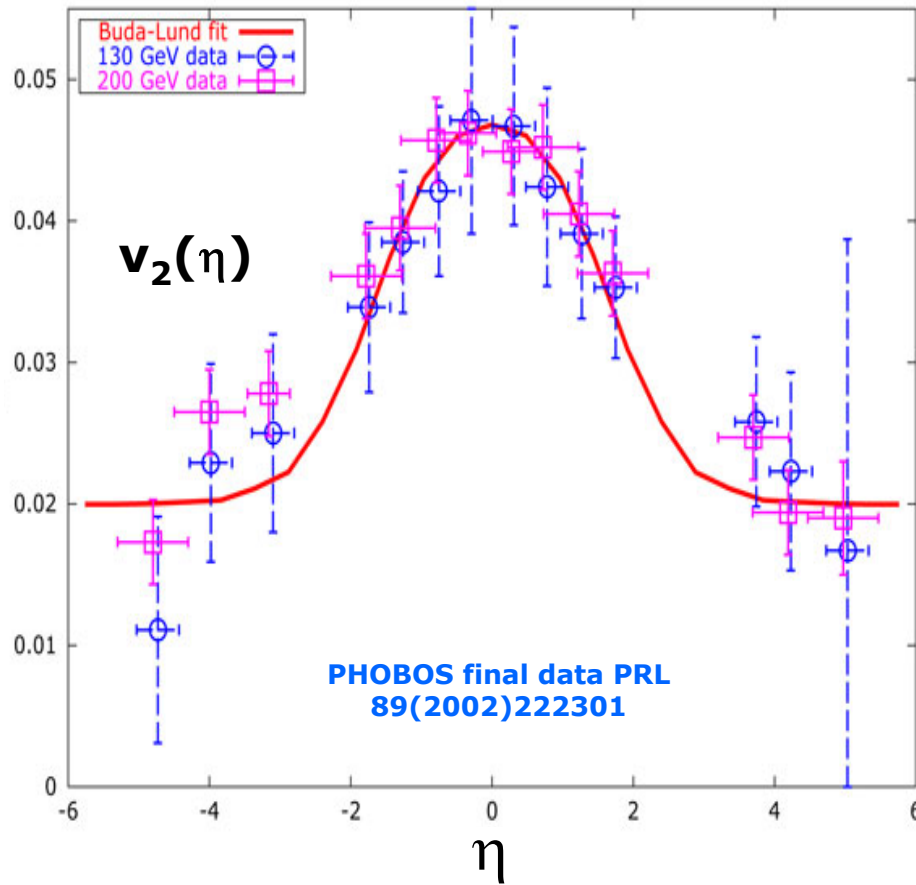
# A solution of the RHIC HBT puzzle

- Our answer: hot center, a fireball heated from inside
- If we **assume** a hot center, the difference between out and side vanishes



- **Confirmed** in fits for the rapidity dependence of the elliptic flow, see poster (**Flow-12**) and nucl-th/0310040
- **Prediction:**  $v_4(p_t)$ ,  $v_1(\eta)$ ,  $v_3(\eta)$

# Confirmation of hot center



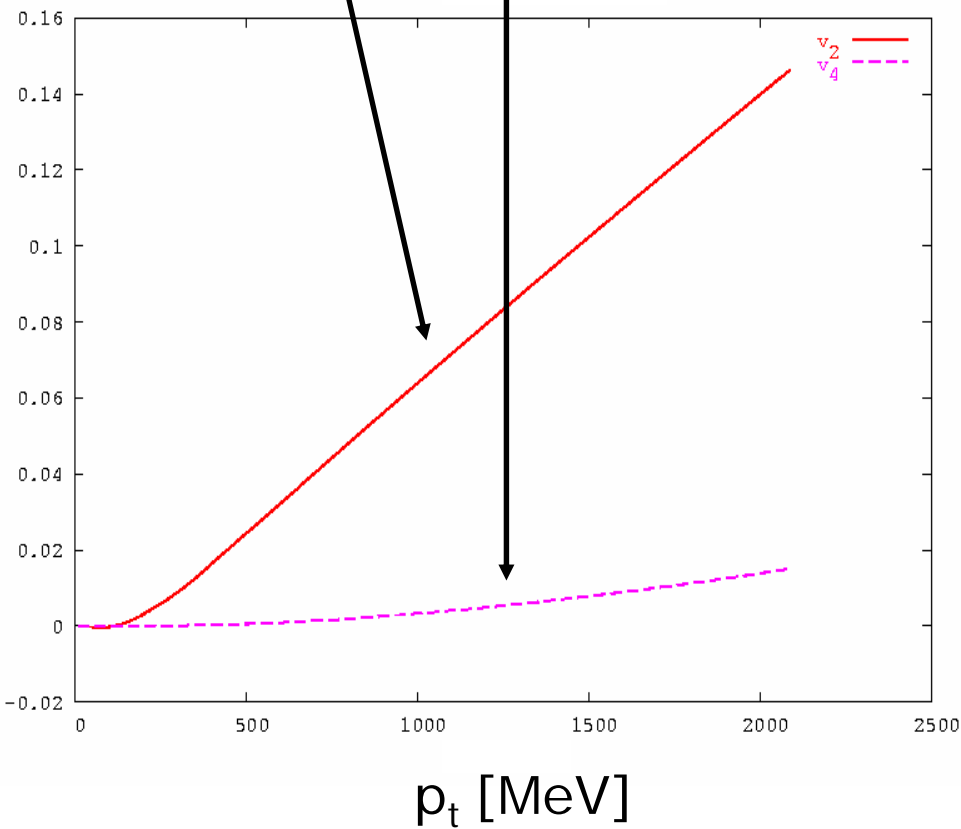
$$T_s = 105 \text{ MeV}, T_0 = 210 \text{ MeV}, \tau_0 = 7 \text{ fm}/c, \vartheta = 0.09$$

$$\dot{X} = 0.57, \dot{Y} = 0.45, \dot{Z} = 2.4$$

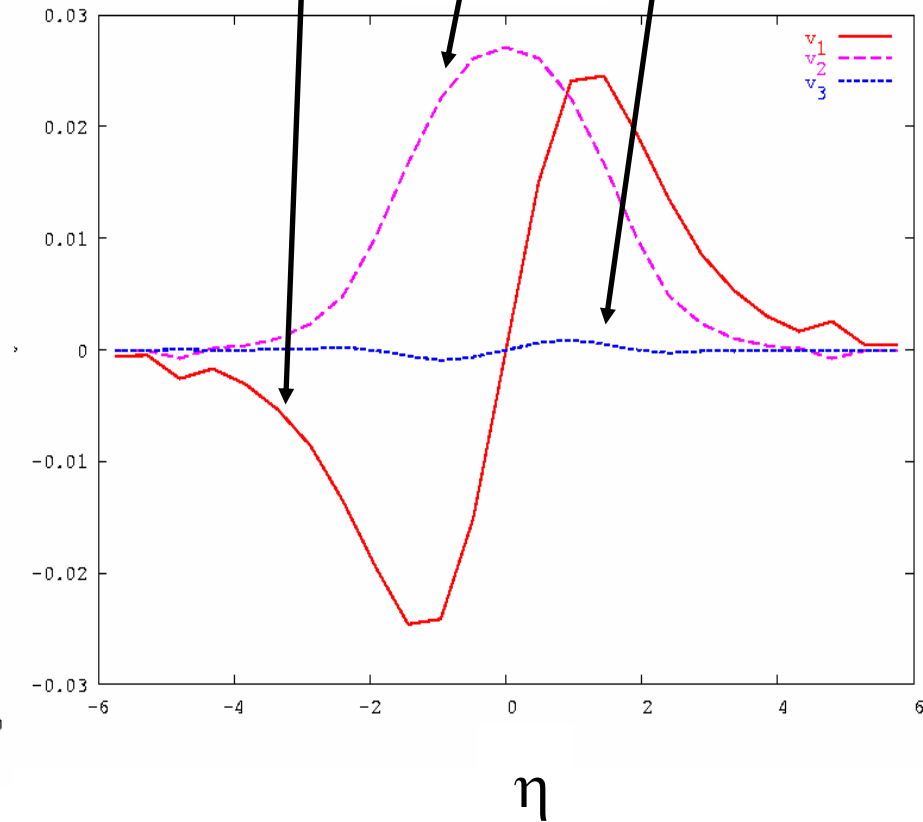
$$X_f = 8.6 \text{ fm}, Y_f = 10.5 \text{ fm}, Z_f = 17.5 \text{ fm}$$

# Prediction

$v_2(p_t), v_4(p_t)$



$v_1(\eta), v_2(\eta), v_3(\eta)$

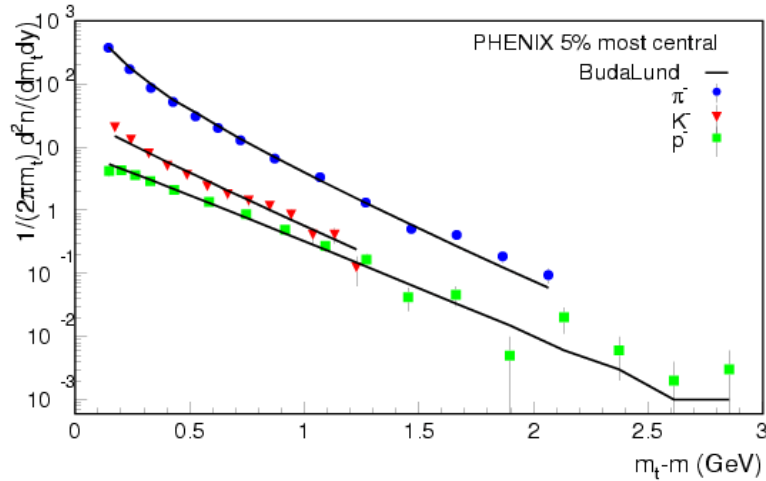




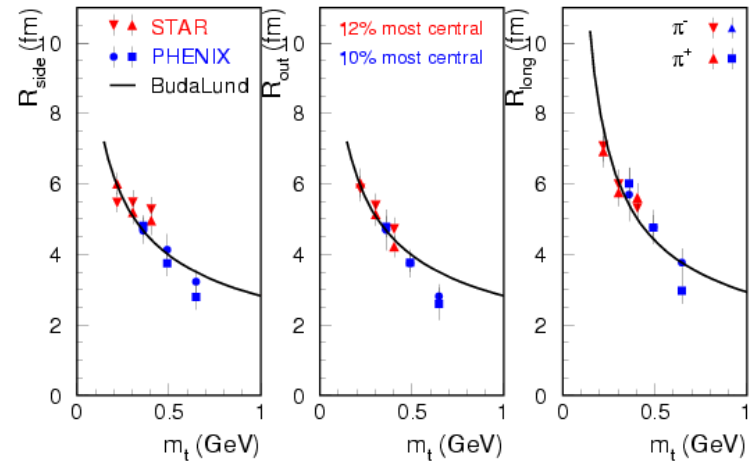
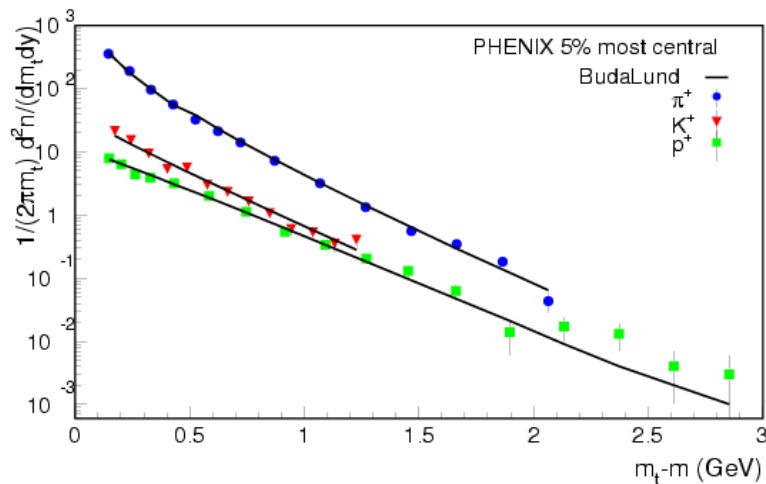
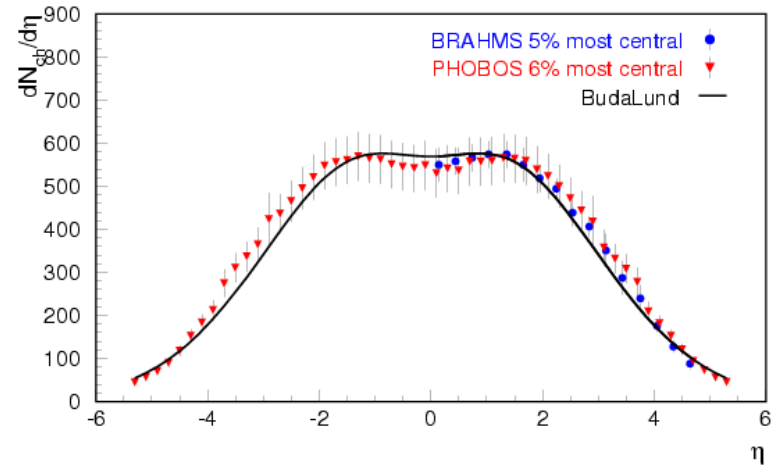
**Thank you for your attention.**

# BudaLund fits to final run-1 RHIC data

BudaLund hydro v1.4, 130 AGeV Au+Au



Buda-Lund v1.4 fits to 130 AGeV Au+Au



# BudaLund fits to final run-1 RHIC data

BL-H source	Buda-Lund v1.0	Buda-Lund v1.4
$T_0$ [MeV]	$202 \pm 13$	$232 \pm 16$
$\langle u_t \rangle$	$1.0 \pm 0.2$	$0.9 \pm 0.06$
$R_G$ [fm]	$9.8 \pm 1.2$	10.0 Fixed
$\tau_0$ [fm/c]	$6.1 \pm 0.3$	$6.1 \pm 0.3$
$\Delta\tau$ [fm/c]	$0.0 \pm 1.5$	$0.0 \pm 1.0$
$\Delta\eta$	2.5 Fixed	$2.3 \pm 0.04$
$T_{\text{surf}}$ [MeV]	$110 \pm 24$	$94 \pm 4$
$T_{\text{eva}}$ [MeV]	$88 \pm 25$	$98 \pm 11$
$\mu_0^{\pi^-}$ [MeV]	$75 \pm 19$	$97 \pm 16$
$\mu_0^{K^-}$ [MeV]	$107 \pm 14$	$89 \pm 27$
$\mu_0^{p^-}$ [MeV]	$305 \pm 41$	$272 \pm 47$
$\mu_0^{\pi^+}$ [MeV]	-	$79 \pm 13$
$\mu_0^{K^+}$ [MeV]	-	$123 \pm 27$
$\mu_0^p$ [MeV]	-	$353 \pm 46$
$\chi^2/\text{NDF}$	74/68 =1.08	226/18 =1.25
CL	28.9 %	1.3 %

V1.0 fits to STAR & PHENIX data at midrapidity,  $\Delta\eta$  fixed

V1.4  $R_G$  fixed as data determine  $\langle u_t \rangle / R_G$  dependence only

# NEXT TO DO

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**Fits to :**

**new NA49 data (with Kaon correlation),**

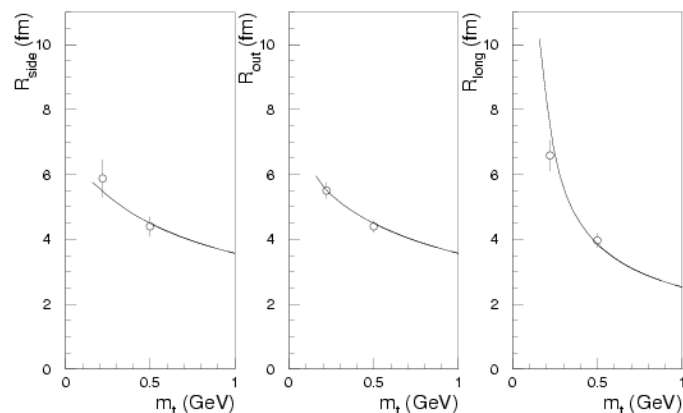
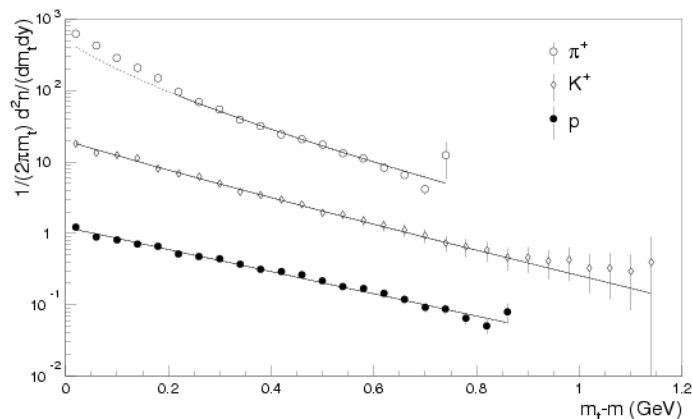
**NA44 data,**

**CERES data**

**at SPS.**

**new run-2 data at RHIC**

# Buda-Lund fits to NA44 Pb+Pb data



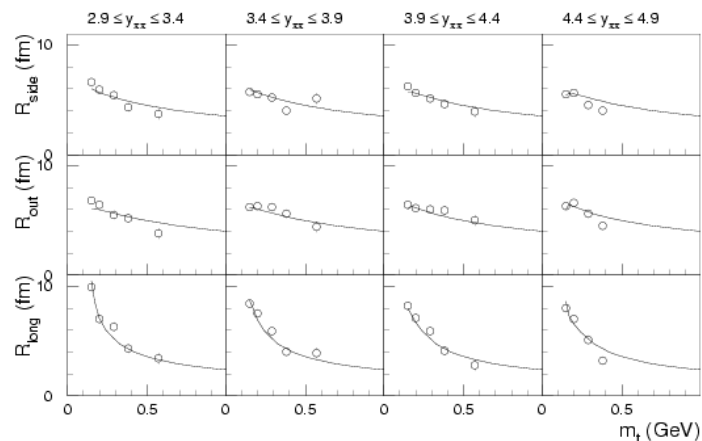
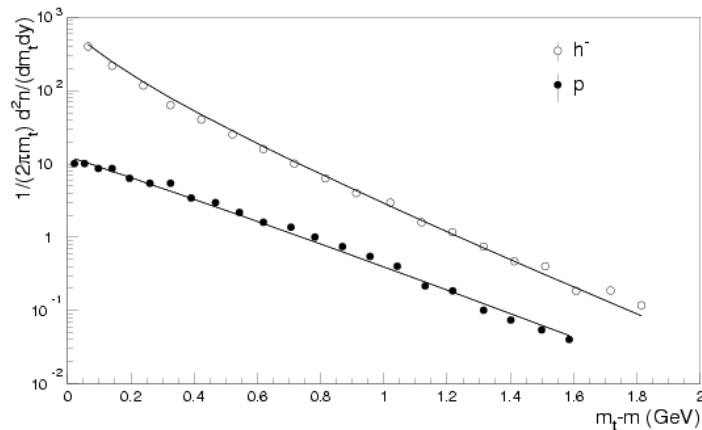
## BL-H source parameters

$T_0$	$145 \pm 3$ MeV
$\langle u_t \rangle$	$0.57 \pm 0.12$
$R_G$	$6.9 \pm 1.1$ fm
$\tau_0$	$6.1 \pm 0.9$ fm/c
$\Delta\tau$	$0.1 \pm 2.2$ fm/c
$\Delta\eta$	$2.4 \pm 1.6$
$\langle \Delta T/T \rangle_r$	$0.08 \pm 0.08$
$\langle \Delta T/T \rangle_t$	$0.87 \pm 0.72$
$\chi^2/\text{NDF}$	$63/71 = 0.89$

**Final data**  
**Absolute normalization,**  
**Boltzmann approx.,**  
 **$\Omega \sim 1$ ,  $\mu_0 = 0$  approx.**

**A. Ster, T. Cs, B. Lörstad, hep-ph/9907338**

# Buda-Lund fits to NA49 Pb+Pb data



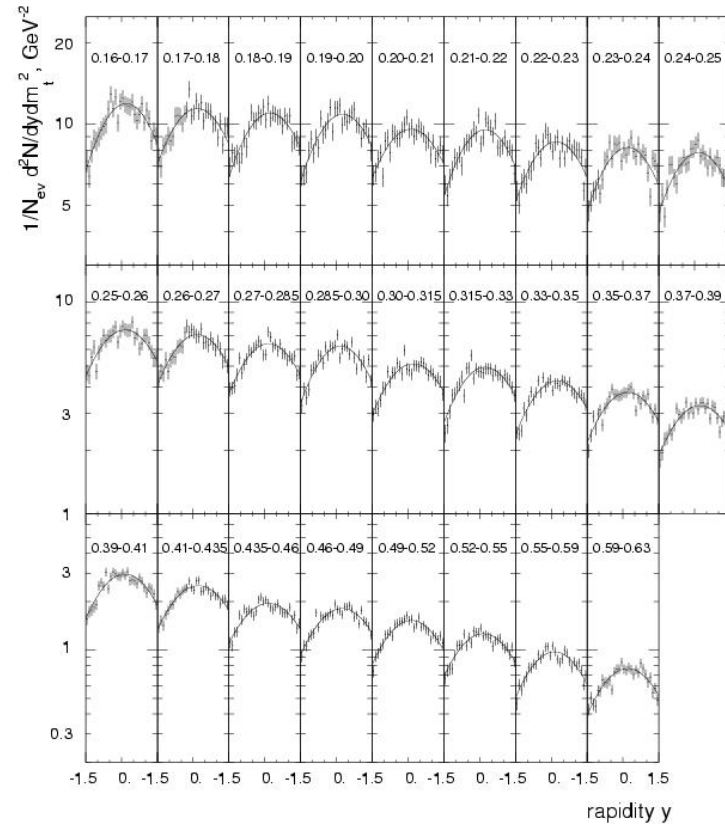
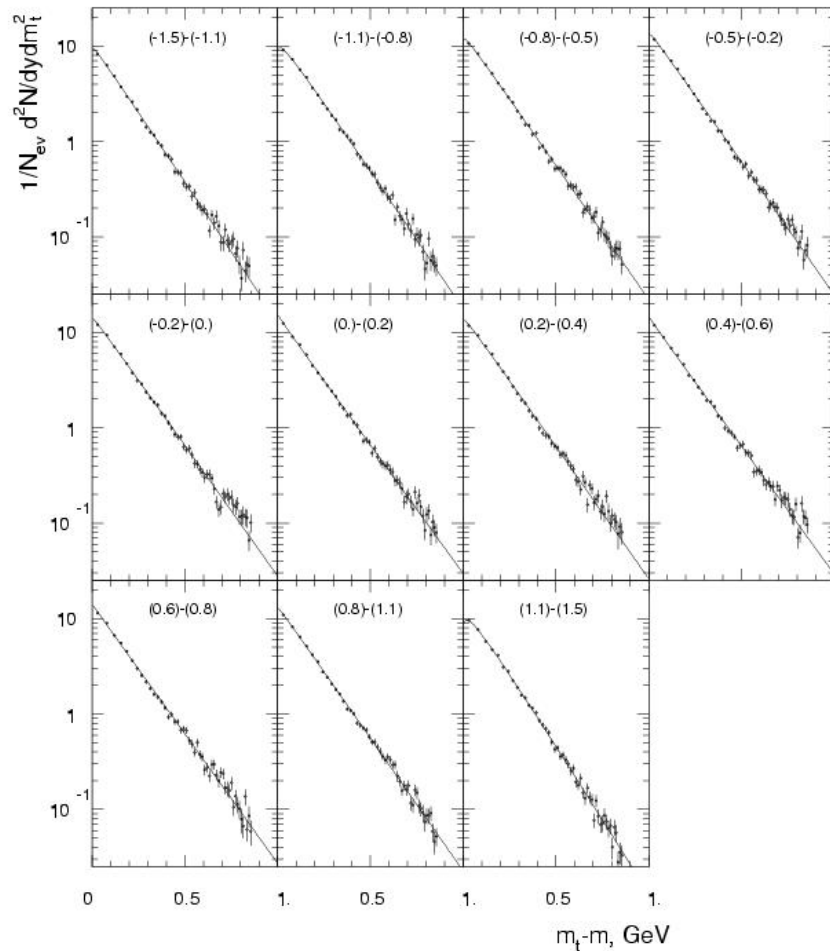
## BL-H source parameters

$T_0$	$134 \pm 3$ MeV
$\langle u_t \rangle$	$0.61 \pm 0.05$
$R_G$	$7.3 \pm 0.3$ fm
$\tau_0$	$6.1 \pm 0.2$ fm/c
$\Delta\tau$	$2.8 \pm 0.4$ fm/c
$\Delta\eta$	$2.1 \pm 0.2$
$\langle \Delta T/T \rangle_r$	$0.07 \pm 0.02$
$\langle \Delta T/T \rangle_t$	$0.16 \pm 0.05$
$\chi^2/\text{NDF}$	$163/98 = 1.66$

**Final data**  
**Absolute normalization,**  
**Boltzmann approx.,**  
 **$\Omega \sim 1$ ,  $\mu_0 = 0$  approx.**

**A. Ster, T. Cs, B. Lörstad, hep-ph/9907338**

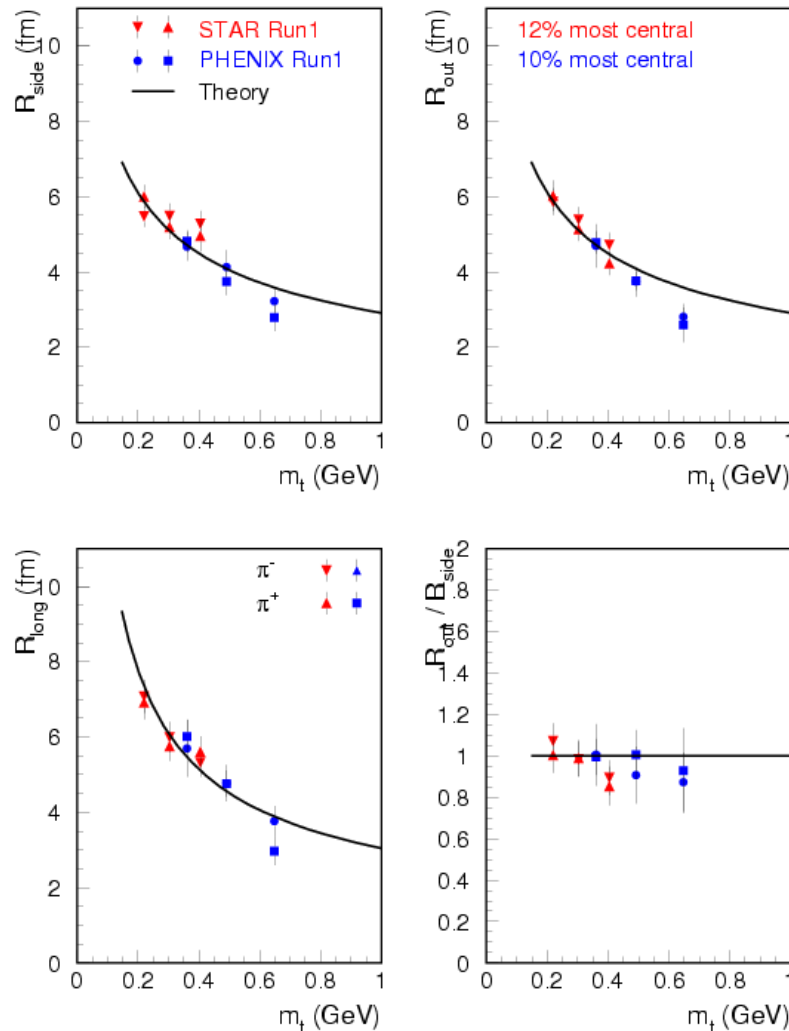
# Buda-Lund fits to NA22 h + p data



N. M. Agababyan et al, EHS/NA22 , PLB 422 (1998) 395  
 T. Csörgő, hep-ph/001233, Heavy Ion Phys. 15 (2002) 1-80

# Comparison of results of models

nucl-th/0207016-1 fits to 130 AGeV Au+Au

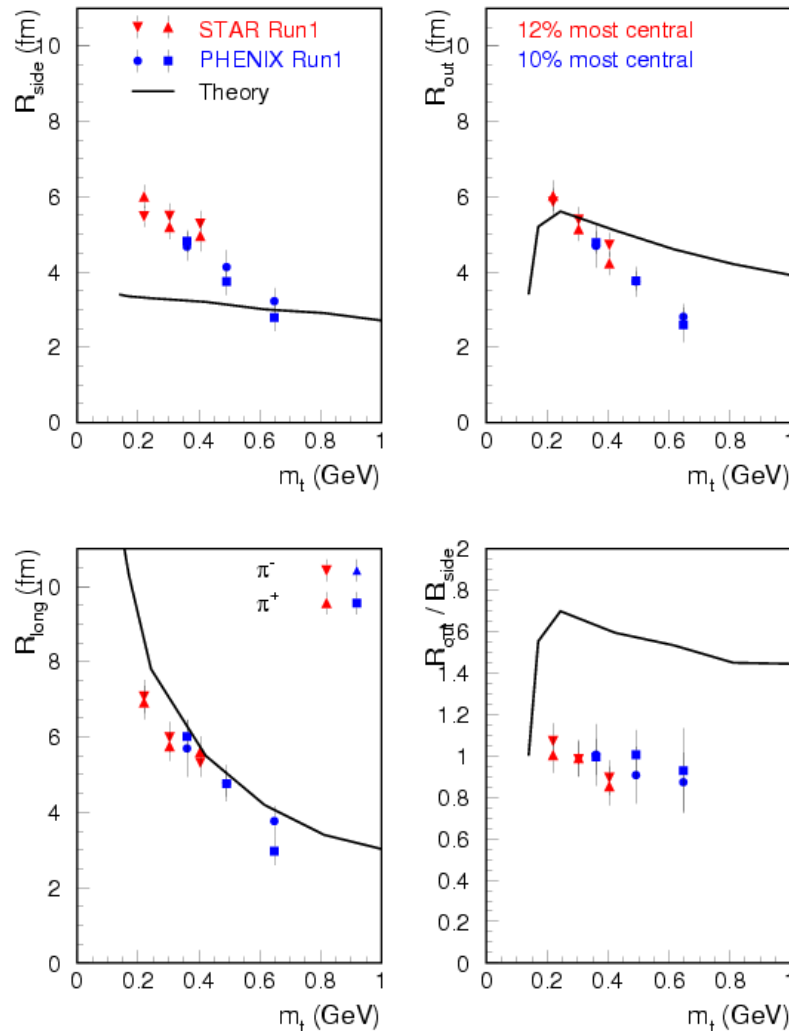


**Acceptable**



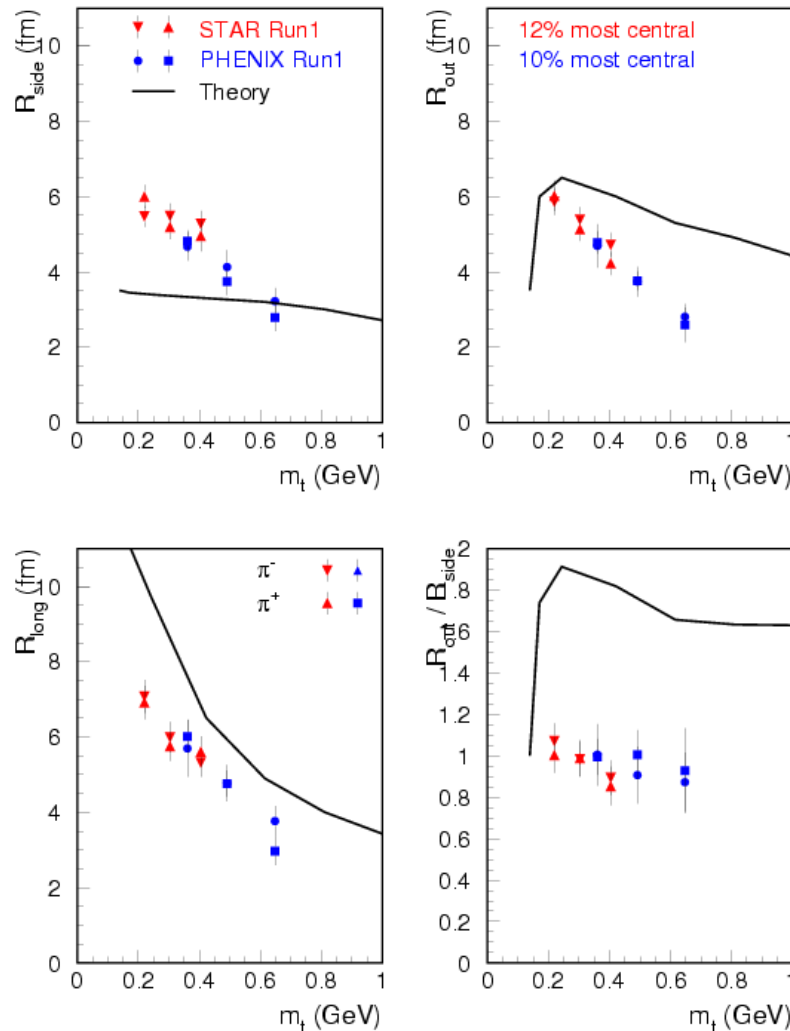
# Comparison of results of models

nucl-th/0208068-1 fits to 130 AGeV Au+Au



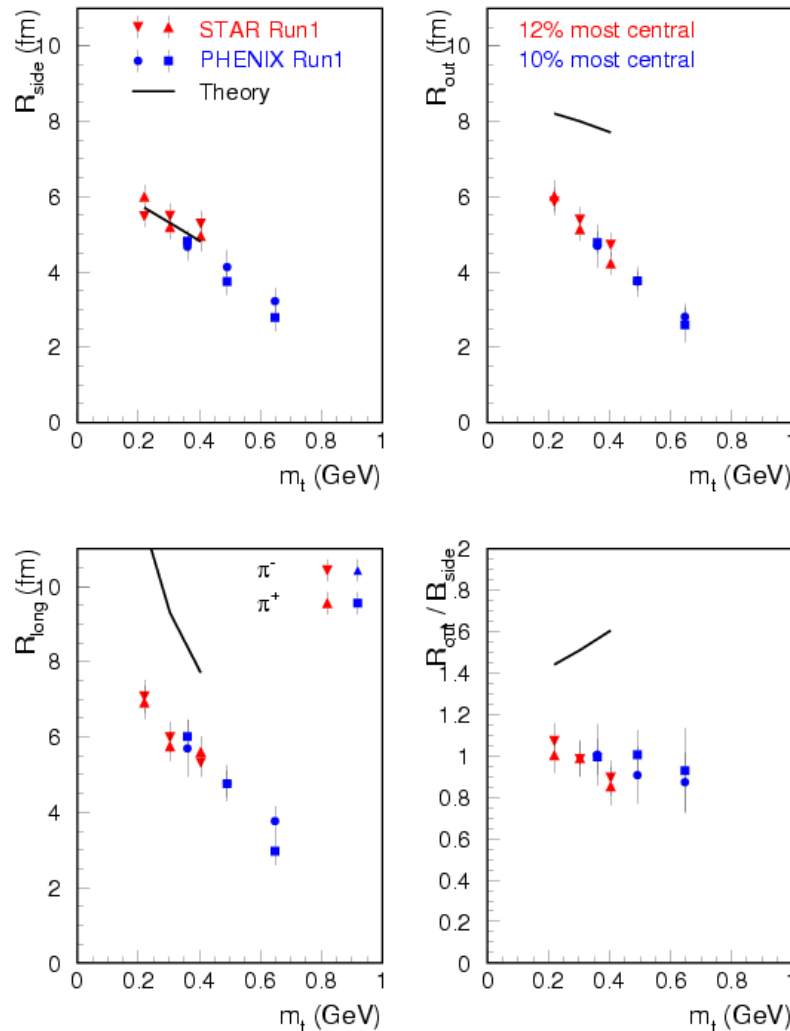
# Comparison of results of models

nucl-th/0208068-2 fits to 130 AGeV Au+Au



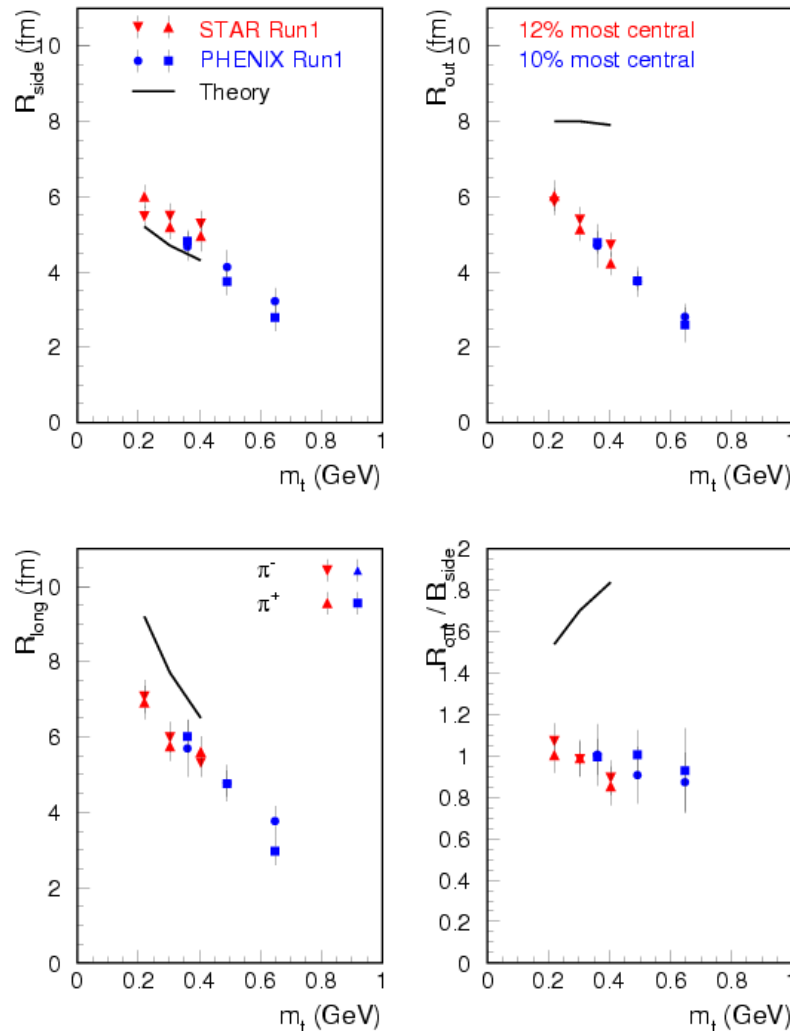
# Comparison of results of models

nucl-th/0209055-1 fits to 130 AGeV Au+Au



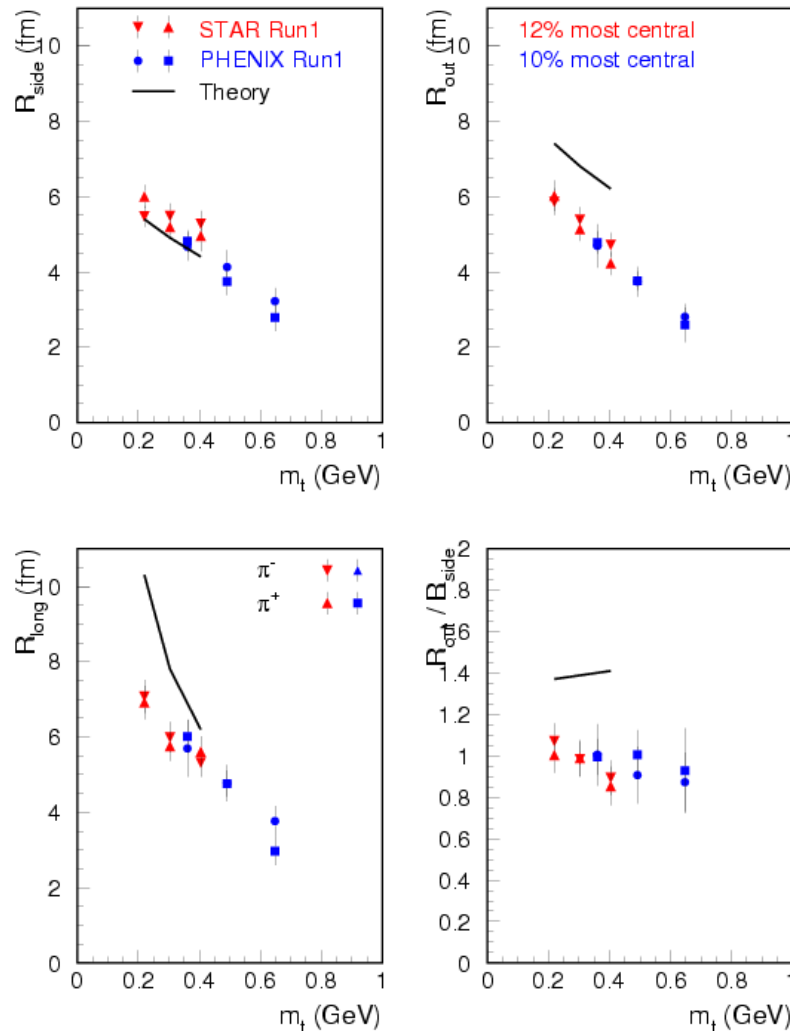
# Comparison of results of models

nucl-th/0209055-2 fits to 130 AGeV Au+Au



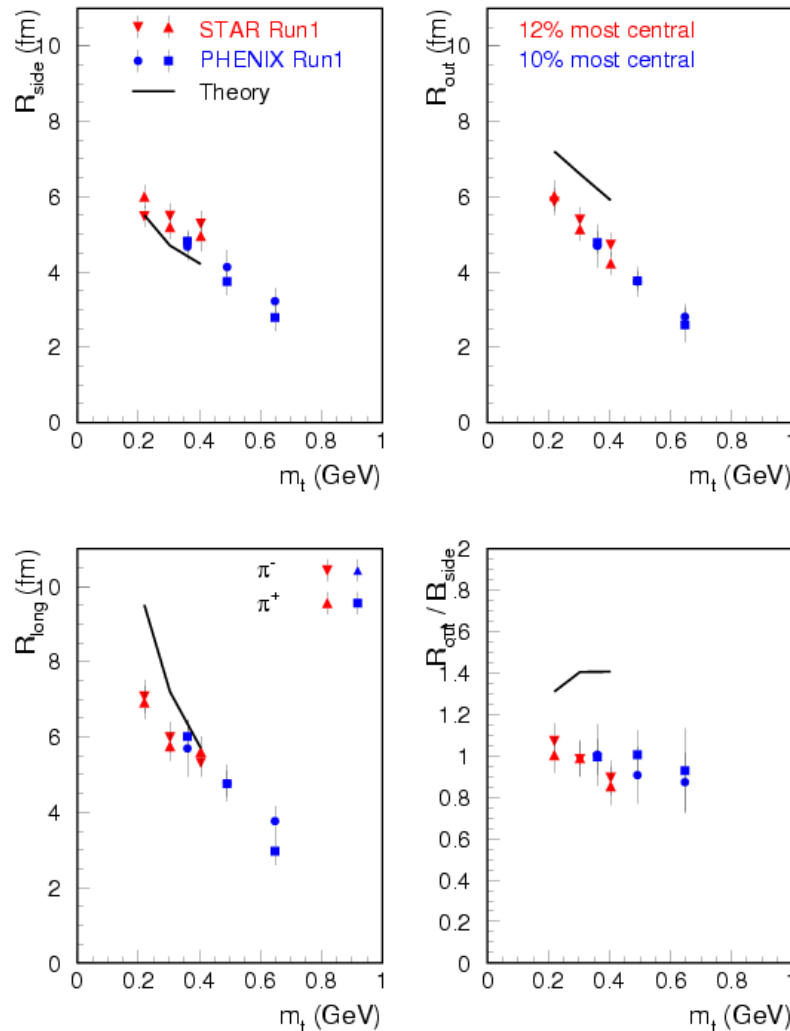
# Comparison of results of models

nucl-th/0209055-3 fits to 130 AGeV Au+Au



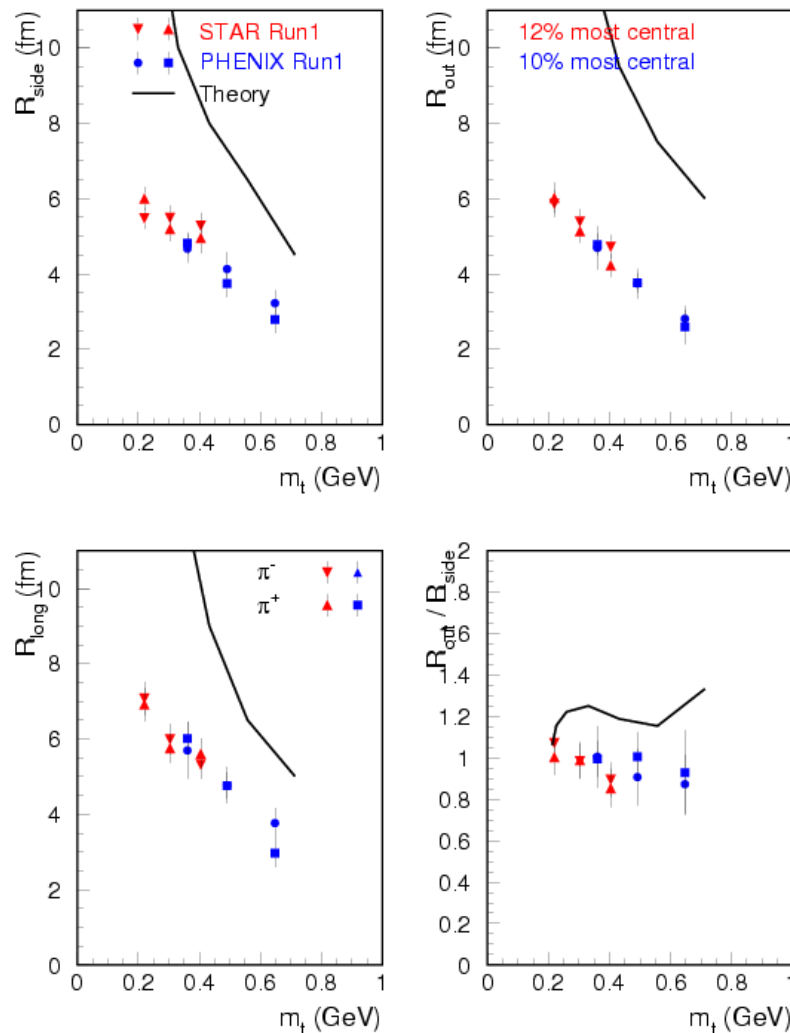
# Comparison of results of models

nucl-th/0209055-4 fits to 130 AGeV Au+Au



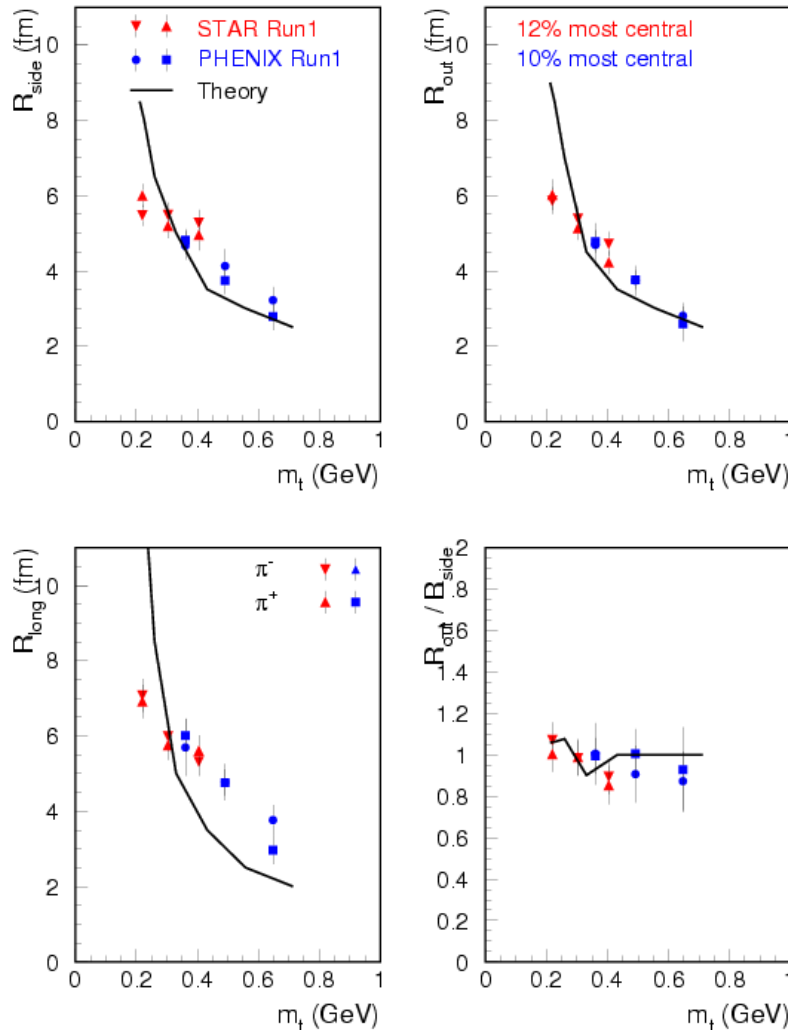
# Comparison of results of models

nucl-th/0204054-1 fits to 130 AGeV Au+Au



# Comparison of results of models

nucl-th/0204054-2 fits to 130 AGeV Au+Au

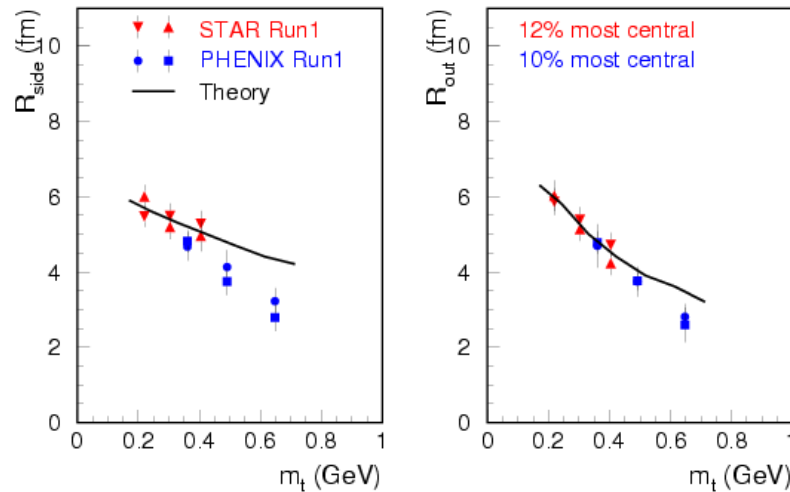


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**Acceptable**

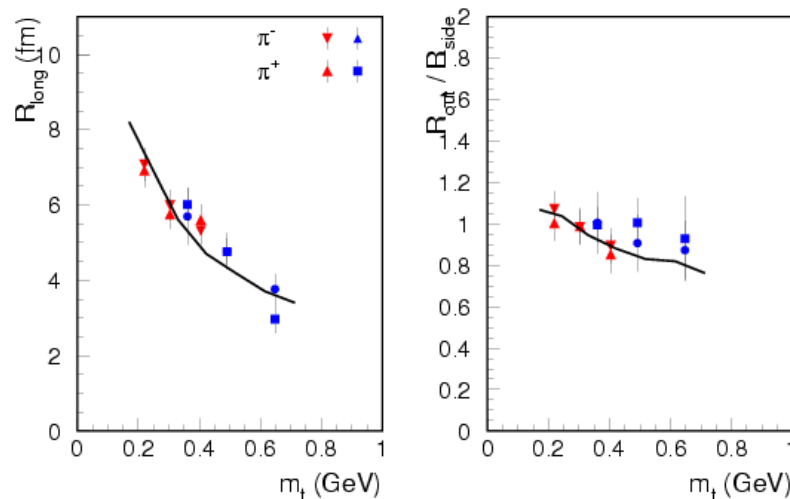


# Comparison of results of models

nucl-ex/0307026-1 fits to 130 AGeV Au+Au

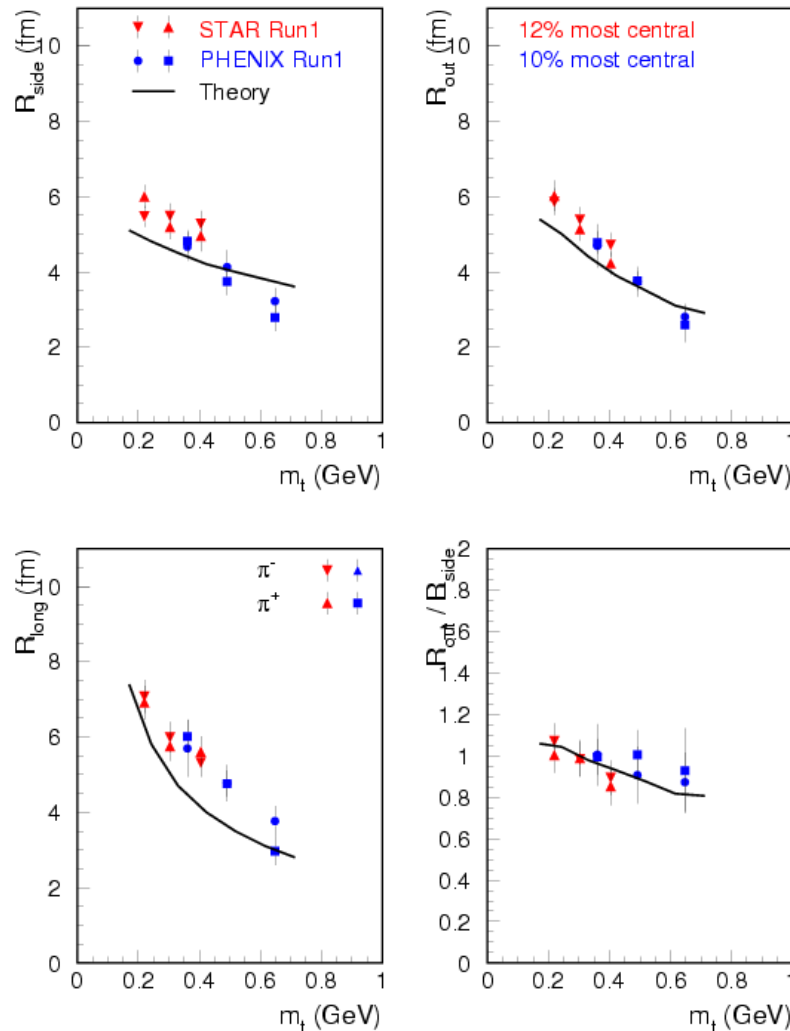


**Acceptable**



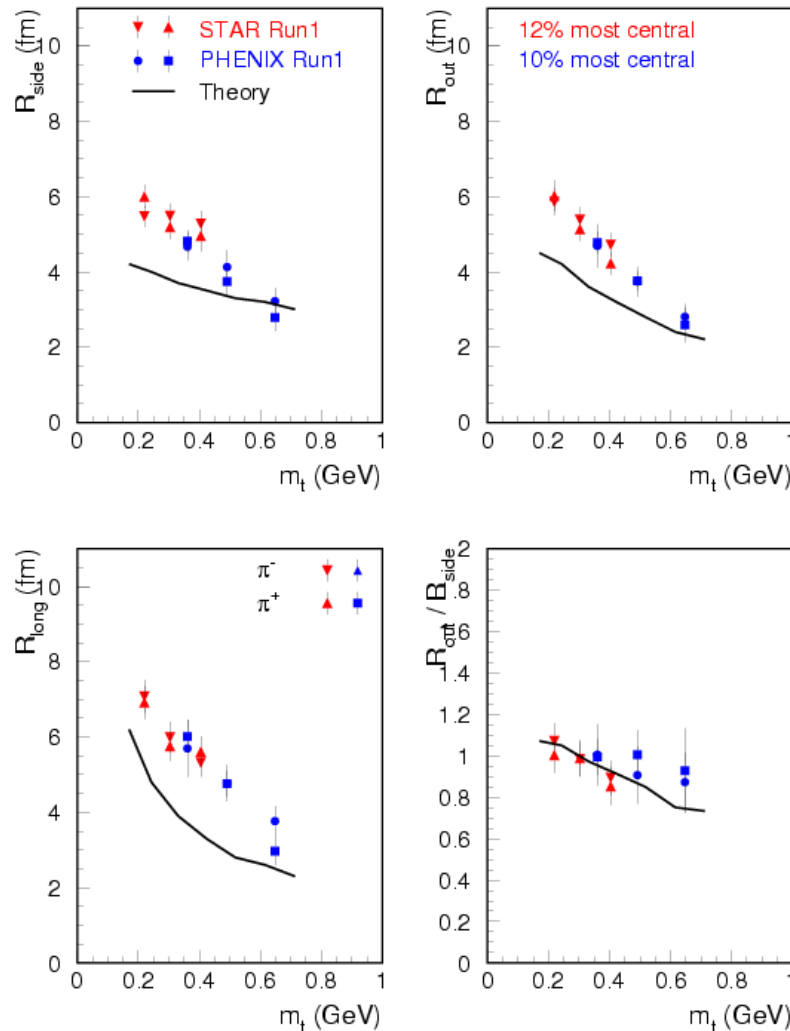
# Comparison of results of models

nucl-ex/0307026-2 fits to 130 AGeV Au+Au



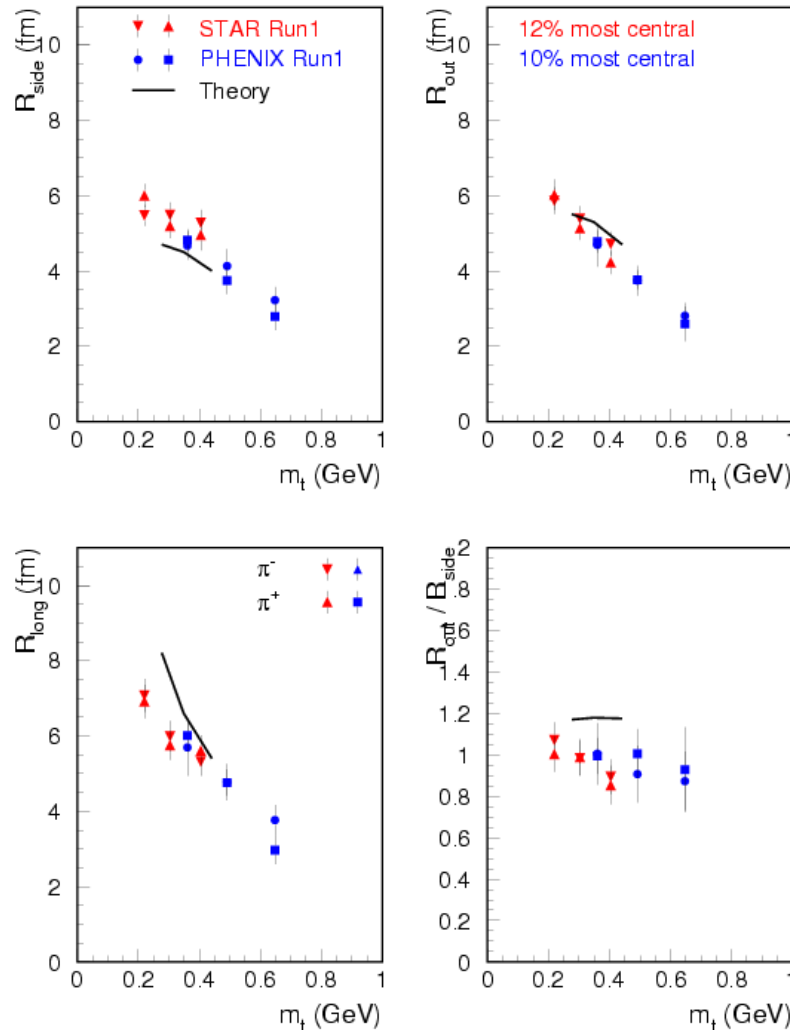
# Comparison of results of models

nucl-ex/0307026-3 fits to 130 AGeV Au+Au



# Comparison of results of models

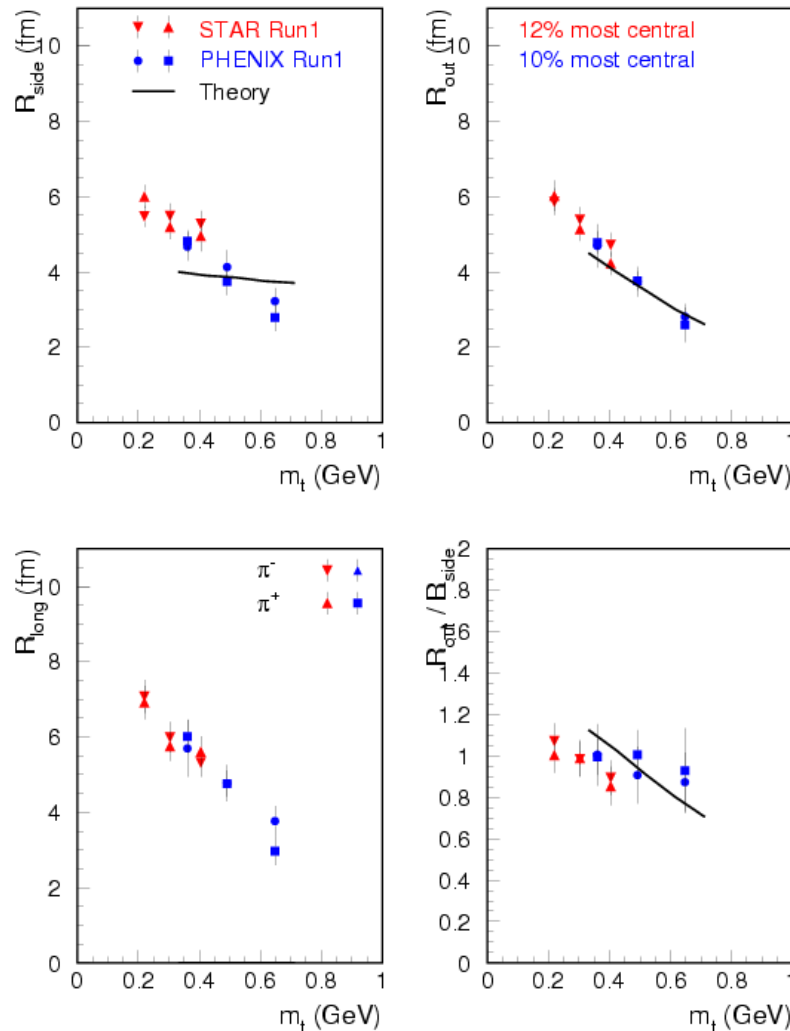
nucl-th/0205053-1 fits to 130 AGeV Au+Au



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Acceptable

# Comparison of results of models

nucl-th/0212053-2 fits to 130 AGeV Au+Au



# Comparison of results of models

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## Models with acceptable results:

nucl-th/0204054

**Multiphase Trasport model (AMPT).**  
Z. Lin, C. M. Ko, S. Pal.

nucl-th/0205053

**Hadron cascade model.** T. Humanic.

nucl-th/0207016

**Buda-Lund hydro model.**  
T.Csörgő. A. Ster,  
Heavy Ion Phys. 17 (2003) 295-312.

nucl-ex/0307026

**Blast wave model.** F. Retière for STAR.

**Not shown here but acceptable:**

nucl-th/0208068

**3D hydro model.** T. Hirano, & T.Tsuda.

hep-ph/0209054

**Cracow (single freeze-out, thermal) model.**  
W.Broniowski, A. Baran, W. Florkowski.