

Hydrodynamic Models of Heavy-Ion Collisions

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Parallel Talks Based on Hydro Jan. 13 •H. Niemi, Photon production from non-equilibrium QGP in heavy-ion collisions •M. Csanad, Indication for quark deconfinement and evidence for a Hubble flow in Au+Au collisions at RHIC Jan. 15 •Y. Nara, CGC, hydrodynamics and the parton energy loss •E. Shurvak, Why does the QGP behaves like a perfect fluid? •U. Heinz, Rapidity dependence of momentum anisotropies in nuclear collisions •D. Teaney, Viscosity and thermalization Tetsufumi Hirano (RBRC) 2

Outline 1. Why hydrodynamics? 2. How hydrodynamics works at RHIC 3. Hybrid models based on hydrodynamics - Information of the inside (jet quenching, EM probe) - Improvement of initial stage 4. Improvement of ideal hydro (viscosity) 5. Summary

1. Why Hydrodynamics?
Once we accept local thermalization ansatz, life becomes very easy
Static
EoS from Lattice QCD
Finite *T*, μ field theory
Critical phenomena

Energy-momentum:

Conserved number: ¿

Dynamic Phenomena in HIC
Expansion, Flow
Space-time evolution of thermodynamic variables

Caveat: Thermalization in HIC is a tough problem like building the Golden Gate Bridge!

•Chiral property of hadron

1. Why Hydrodynamics (contd.) Space-time evolution of energy density in $sqrt(s_{NN})=200$ GeV Au+Au collision at b=7.2fm

Animation is here in the presentation. If you need, please ask me (hirano@bnl.gov).

A full 3D hydrodynamic simulation with a CGC initial condition Talk by Y.Nara

Hydrodynamics provides us a very intuitive and simple description of relativistic heavy ion collisions.

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2. How Hydrodynamics Works at RHIC

Elliptic flow (J.-Y.Ollitrault ('92)) How does the system respond to initial spatial anisotropy?

Dense or dilute? If dense, thermalization? If thermalized, EoS?

 $\frac{dN}{p_T dp_T dy d\phi} = \frac{1}{2\pi} \frac{dN}{p_T dp_T dy} (1 + 2v_1 \cos(\phi) + 2v_2 \cos(2\phi) + \cdots)$ A.Poskanzer & S.Voloshin ('98) $= \frac{\int d\phi \cos(2\phi) \frac{dN}{p_T dp_T dy d\phi}}{\int d\phi \frac{dN}{p_T dp_T dy d\phi}} = \langle \cos(2\phi) \rangle$

Elliptic Flow of Charged Particles

P.Kolb et al.('01)





rets<mark>ufumi Hirano (RBRC)</mark>

Roughly speaking, ideal hydro gives a good description $b \lesssim 5$ fm, $p_T \lesssim 1-1.5 \, \, {\rm GeV}/c$, $\mid \eta \mid \lesssim 1-2$ For improvement of models, talk by U.Heinz

More on Elliptic Flow

PHENIX, PRL91('03)182301.



Hydro: P.Huovinen et al. ('01)

See recent excellent reviews,

P.Huovinen (QM2002), nucl-th/0305064; P.Kolb and U.Heinz, nucl-th/0305084; E.Shuryak, hep-ph/0312227, today's talk. Tetsufumi Hirano (RBRC)

STAR, PRC66('02)034904



(Note: Hydro+RQMD gives a better description. D.Teaney et al.('01)) Ideal hydro seems to give a good description at RHIC

What's next?

1. <u>Making the most use of hydro</u> <u>models to study the RHIC physics</u> <u>2. Checking how robust the</u> <u>current results are when</u> <u>hydro models are improved</u>

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3.1 Information inside fluids



Jet quenching is a manifestation of interaction between in a and partons (Talks by G.Moore and I.Vitev

For quantitative analysis, the information about the space into cool in the cool into cool in the

is indispensable.

3.1.1 Hydro as a Tool to Analyze Jet Quenching

Jet quenching analysis taking account of (2+1)D hydro results (M.Gyulassy et al.('02))

Hydro+Jet model (T.H. & Y.Nara ('02)) GLV 1st order formula (M.Gyulassy et al.('00)) ∧ E

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ho\left(au, \mathbf{x}(au)
ight) \cap \left[rac{-2m}{2} u_{\mu}
ight]$

Parton density $\rho(x)$ taken from full 3D hydro simulations Animation is here in the presentation. If you need, please ask me hirano@bnl.gov

Movie and data of $\rho(x)$ are available at http://quark.phy.bnl.gov/~hirano/



Consequence from hadron species dependent p_{T,cross} Hydro+Jet Recombination



"Scaling v₂" Interplay between soft and hard? Recombination mechanism?

3.1.2 Hydro as a Tool to Analyze Electromagnetic Radiation

Thermal photon is a penetratingprobe of QGP(E.Shurvak(?78))

• Production rate (Number per unit space-time volume)

$$E\frac{dR_{\gamma}}{d^3p}(T,\mu) = \frac{-1}{(2\pi)^3}n(T,\mu)\operatorname{Im}\Pi^R{}_{\mu}$$

H.A.Weldon ('83), L.McLerran & T.Toimela ('84) C.Gale & J.Kapusta ('91) Talk by G.Moore

Invariant spectrum of photons

 $E\frac{dN\gamma}{d^3p} = \sum_i \Delta V_i \tilde{E} \frac{dR\gamma}{d^3 \tilde{p}} (T(x_i), \mu(x_i), \tilde{E} = p_\mu u^\mu(x_i))$

D.K.Srivastava & B.Sinha('94), J.Sollfrank et al.('97), J.Alam et al.('01) and a lot of work

Importance of temperature profile



 T, μ

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Novel Temperature Evolution



Caveat: one has to take account of fugacity λ in calculating EM spectra. •QGP phase: $\lambda < 1$ •Hadron phase: $\lambda > 1$ $n(T, \lambda) \sim \lambda n_0(T)$ Compensation between T and λ ? Talk by H.Niemi



3.2 Improvement of Initial Condition -Toward an unified model in HIC-

$e(au_0,x), n_B(au_0,x), u^{\mu}(au_0,x)$

Group	Hydro	Initial condition
M.Gyulassy et al.	SHASTA (2+1D, Bjorken)	HIJING, event-by-event
C.Nonaka et al.	Lagrangian hydro (full 3D)	URASiMA, event average
B.Schlei et al.	HYLANDER (2+1D)	VNI, event average
C.E.Aguiar et al.	SPheRIO (full 3D)	NeXus, event-by-event
L.P.Csernai et al.	Particle-in-cell (full 3D)	String ropes, flux tubes, classical YM
K.Eskola et al.	SHASTA (2+1D, Bjorken)	pQCD + final state saturation
T.H. & Y.Nara	τ – η coordinate (full 3D)	CGC, $\phi(k_T^2, x)$ a la Kharzeev & Levin

*Smoothed Particle hydrodynamical evolution of Relativistic heavy IOn collisions (Sao Paulo & Rio de Janeiro)

3.2.1 SPheRIO*

Main features:

C.E.Aguiar, R.Andrade, F.Grassi, Y.Hama, T.Kodama, T.Osada, O.Socolowski Jr.... Poster by F.Grassi

- "Particle" method (a kind of Lagrangian hydro)
- Numerical cost cheaper than conventional finite grids method (Even in 3+1 D, any geometry)
- Event-by-event physics (NeXus + SPheRIO=NeXSPheRIO) (NeXus: parton based Gribov-Regge theory)

SPheRIO

Conventional approach



M.Gyulassy et al.('97)

Initial Conditions in NeXSPheRIO Energy density in the transverse plane (z=0)



Results from NeXSPheRIO

Pb+Pb 17.3A GeV Centrality: 0 – 5 % with fluctuation (50 events) without fluctuation 250 H $\pi^{-} + K^{-} + \overline{p}$ 200 dN/dy 150 100 p – p 50 V 10000Centrality: 14 - 23 %



Effect of *initial energy density* fluctuation (simple EoS case): $\langle \cdots \rangle = \frac{1}{N} \sum$:(event average) $\langle S \rangle_{\text{final}} \sim \langle S \rangle_{\text{initial}}$ Negative! Multiplicity is reduced by ~10%! p_T slope is not affected largely. \rightarrow v₂(p_T) and its fluctuation? Now the hydro simulation becomes close to experimental situations like event-generators!



CGC+Hydro+Jet Model (contd.)



Initial condition of energy density from CGC Au+Au 200AGeV b=7.2fm, 70 =0.6fm



Results from CGC+hydro+jet

- CGC+hydro

PHOBOS



CGC initial condition works very well! (Energy, rapidity, centrality dependences)



For details, talk by Y.Nara

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4. Viscosity

Talks by E.Shuryak and D.Teaney

Change not only the equations of motion but the local thermal distribution function A.Dumitru('02), D.Teaney('03)

Tetsufumi Hirano (RBRC)

• Blast wave model + dist. fn. with viscous correction 1st order correction to dist. fn.:

 $\delta f \propto \frac{1}{T^2} f_0 (1+f_0) p^{\mu} p^{\nu} X_{\mu\nu}$

s :Sound attenuation length

 $X_{\mu\nu}$: Tensor part of thermodynamic force Reynolds number in Bjorken flow

 $R^{-1} \approx \Gamma_s / \tau$

Nearly ideal hydro !?





5. Summary

Hydrodynamics is one of the valuable tools at RHIC energies

 Open our mind / Hydrodynamics can be used even for "high p_T physics in HIC".

- Jet tomography
- EM probe

...

- $(J/\Psi \text{ suppression})$

- Keep in mind !
 How robust is the current agreement of hydro?:
 - Chemical non-eq.?
 - Initial fluctuation?
 - Viscosity?
 - Thermalization?
 - EoS?
 - (Freeze-out?)