

HBT Measurements “at RHIC”

Mark D. Baker

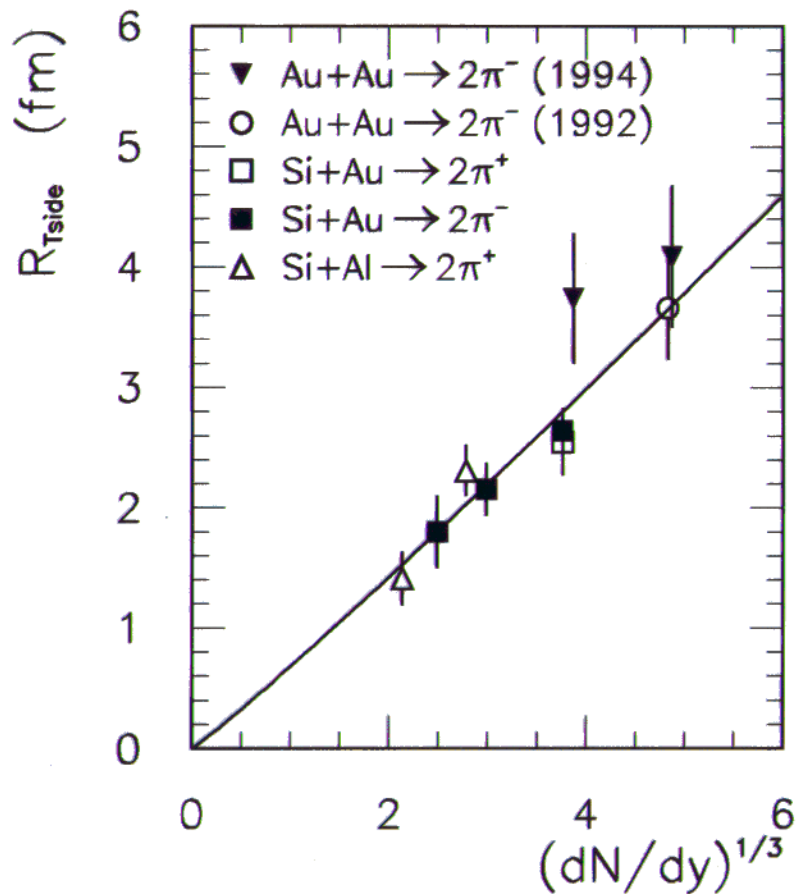
Brookhaven National Laboratory

- Current Data (AGS/SPS)
 - System, rapidity dependence
 - Dynamics: k_T dependence
 - A puzzle (beam E dependence)
- Outlook (RHIC)
- Conclusions

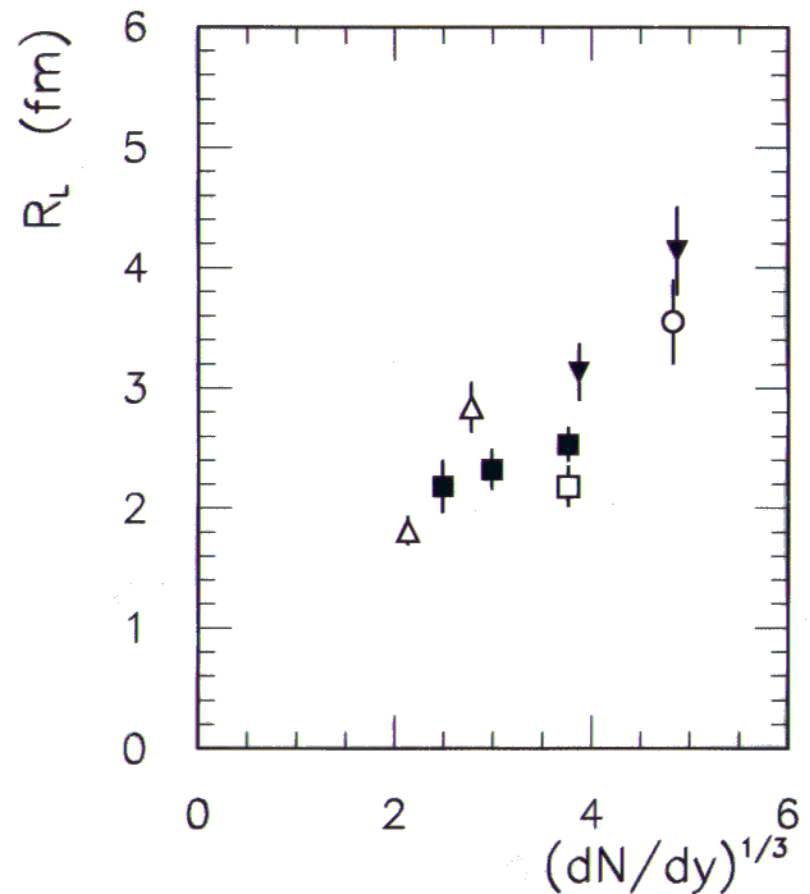
AGS Centrality Dependence (radius)

E866 (Baker, Soltz et al.) Preliminary 11-15 AGeV/c

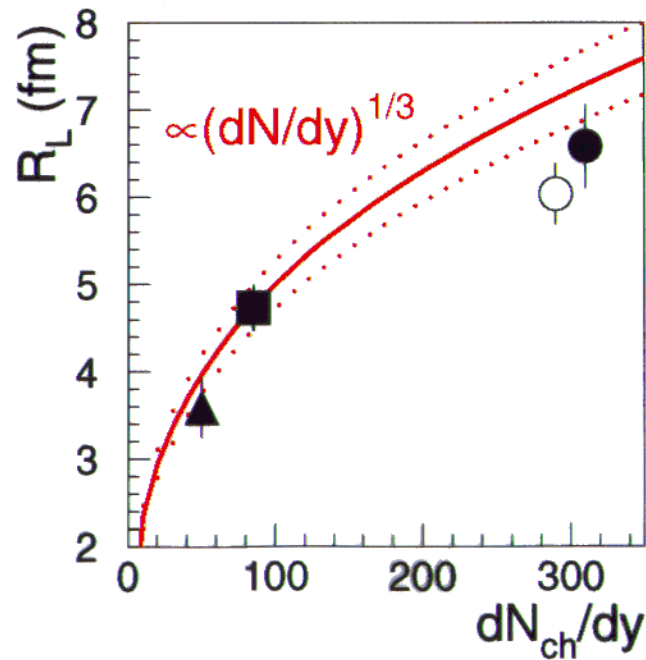
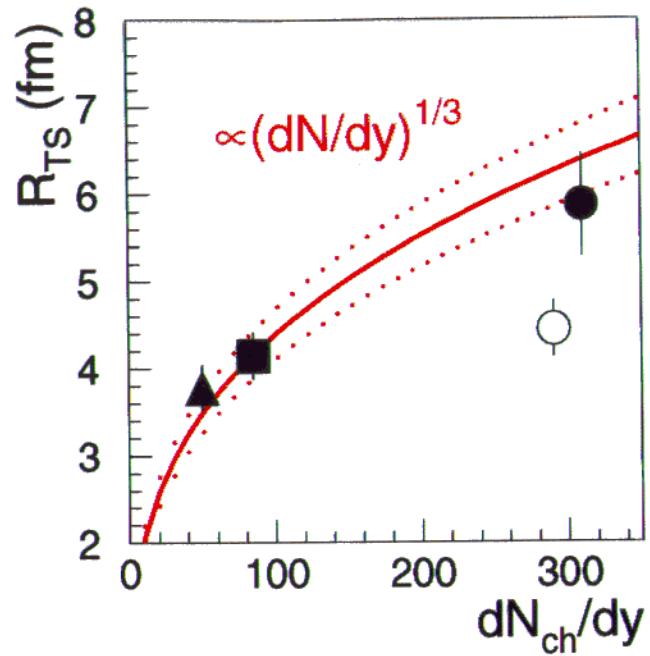
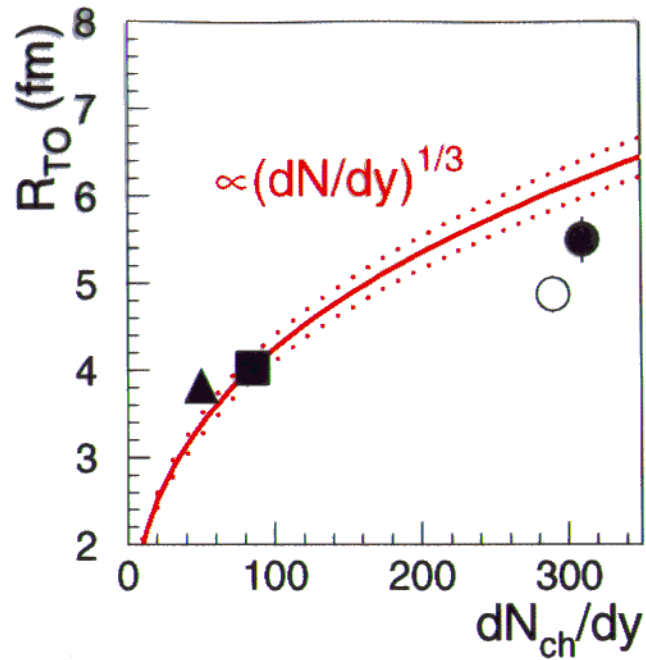
Geometric Radius (transverse)



Longitudinal Length param.



$$R_{Ts} = (0.68 \pm 0.22) \cdot (dN/dy)^{0.355 \pm 0.085}$$

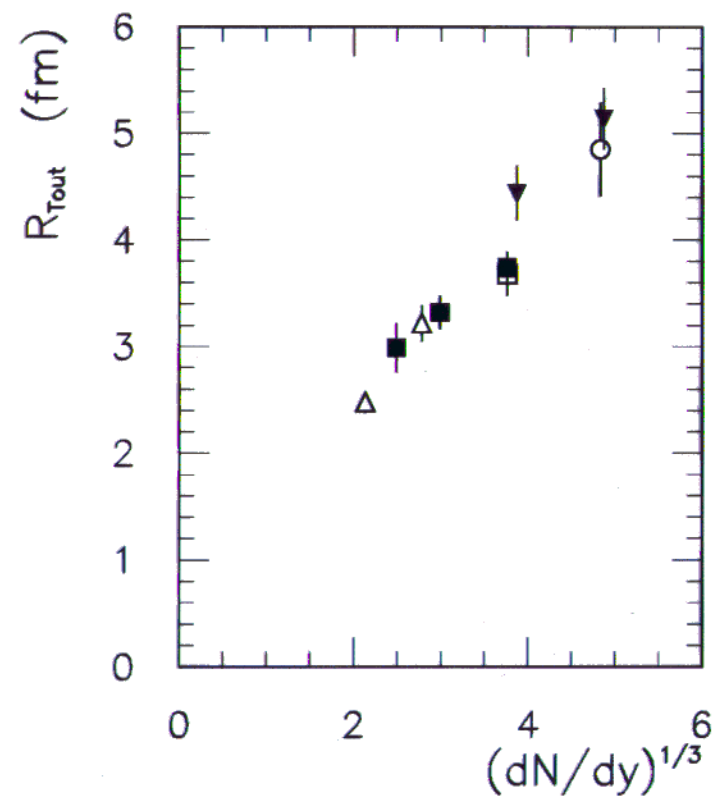
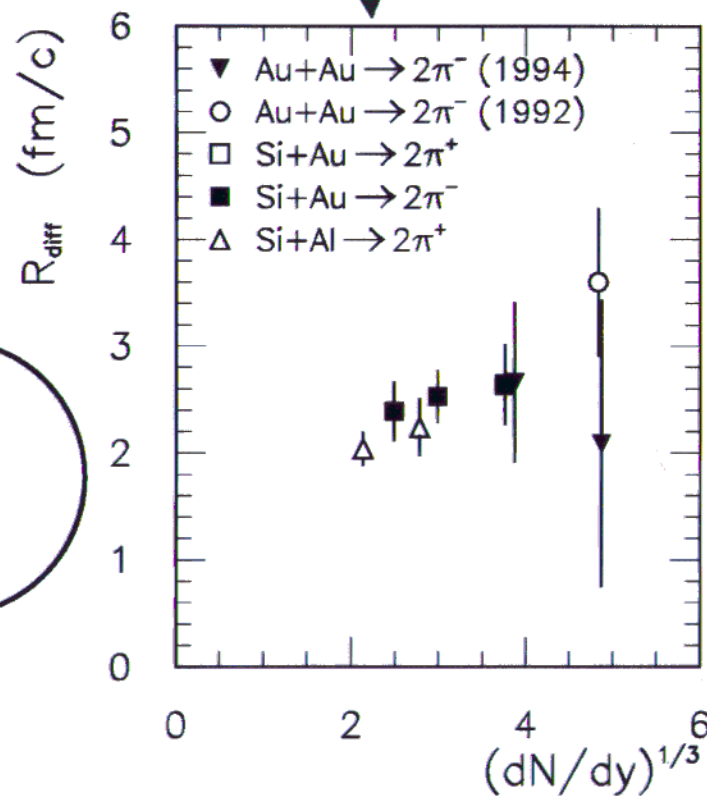
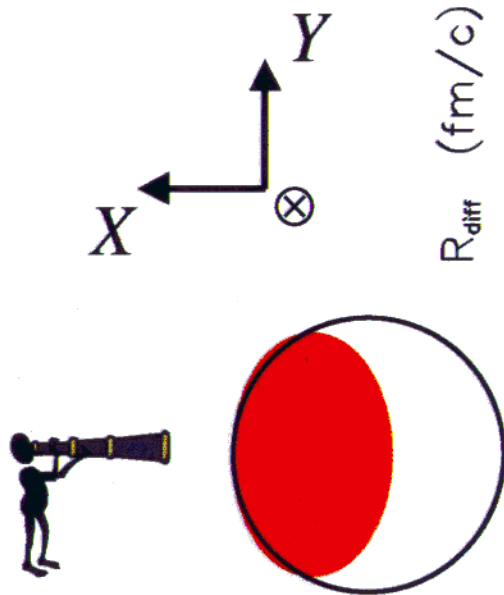


A+A Semi-Central *N*A44

- ▲ S+S 200A GeV $2\pi^+$
- S+Pb 200A GeV $2\pi^+$
- Pb+Pb 158A GeV $2\pi^+$
- Pb+Pb 158A GeV $2\pi^-$

Centrality Dependence (lifetime)

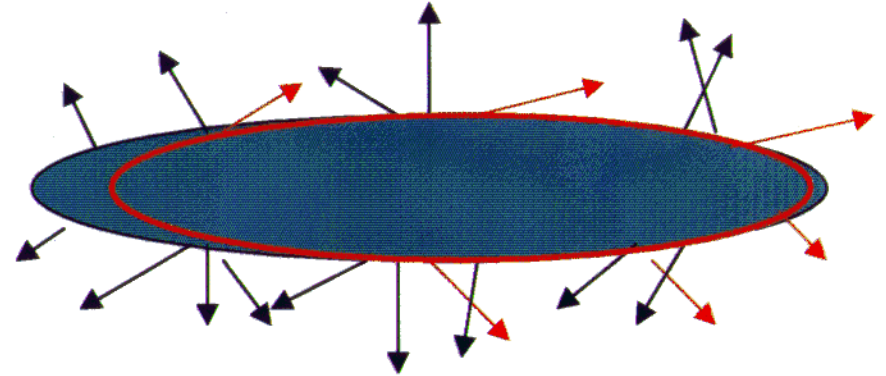
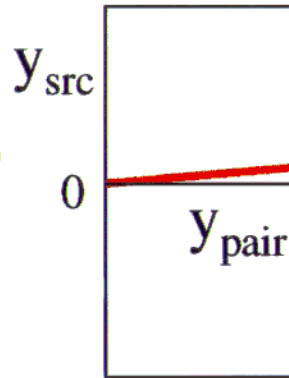
$$R_{diff}^2 = \beta_{\perp}^2 (\Delta\tau)^2 + \underline{(X^2 - Y^2)} - 2\beta_{\perp} \sigma_{xt}^2$$



YKP Source Velocity

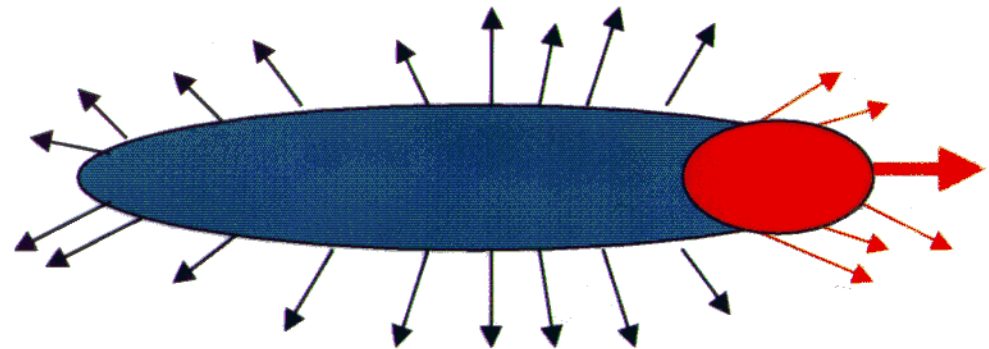
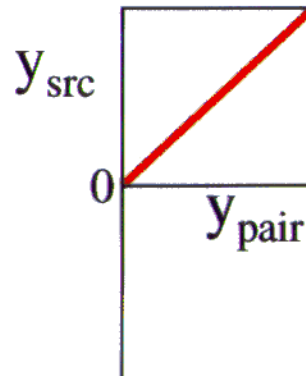
Static Source:

$Y_{\text{source}} \sim 0$ (mid.-rap.)



Scaling Source:

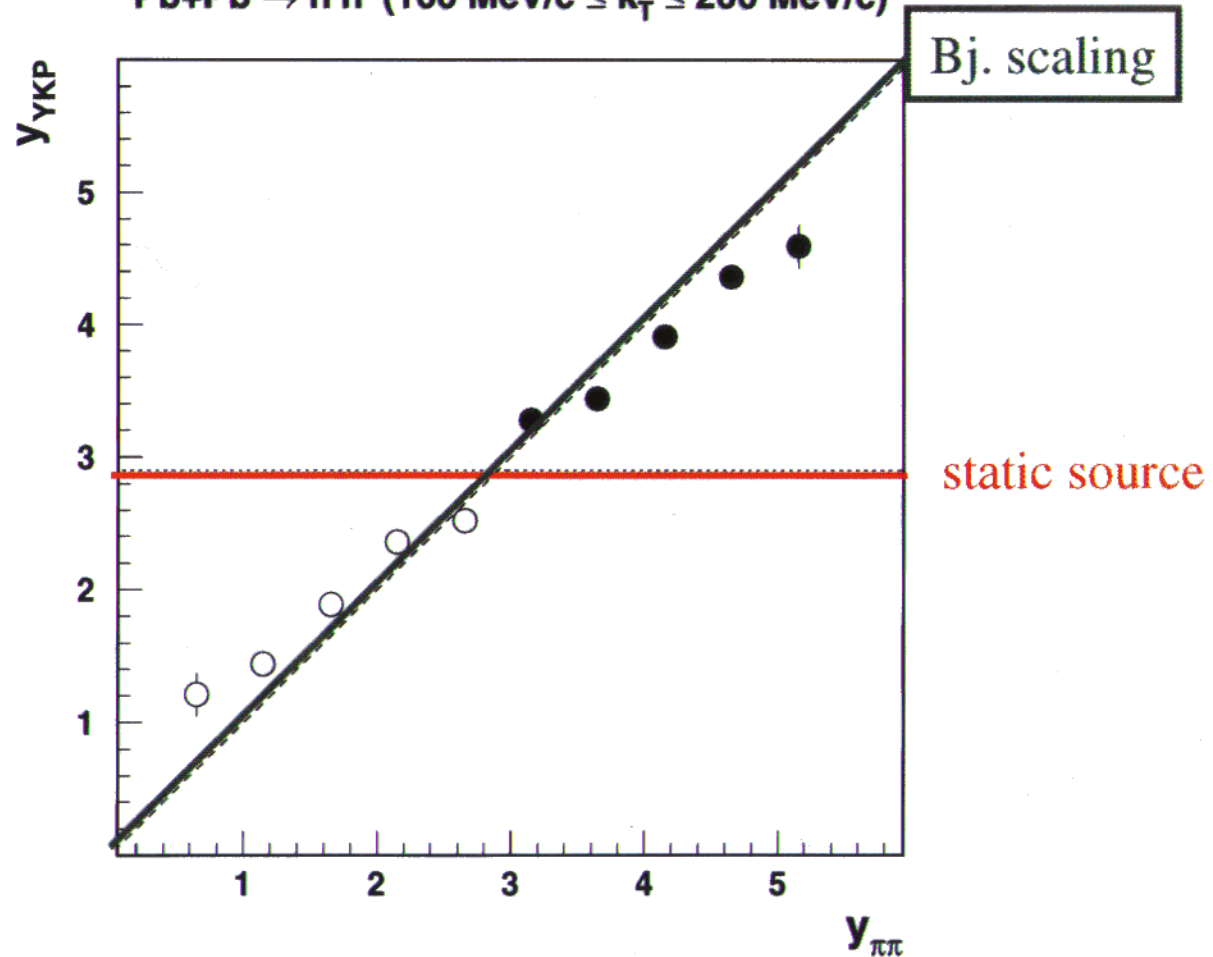
$Y_{\text{source}} = Y_{\text{pair}}$



Source Rapidity (SPS)

NA49 Preliminary(?)

Pb+Pb \rightarrow h \bar{h} ($100 \text{ MeV}/c \leq k_T \leq 200 \text{ MeV}/c$)

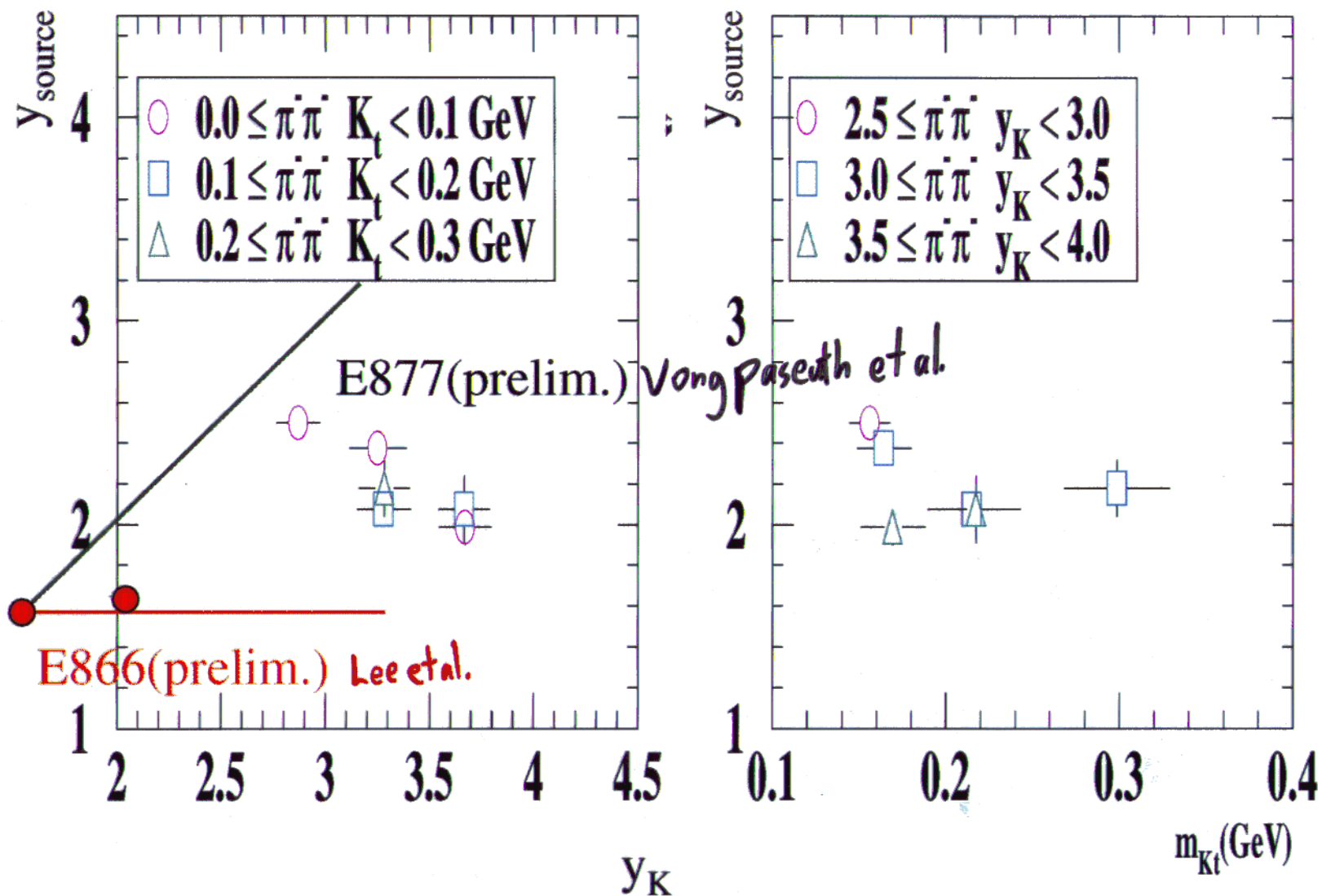


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HBT Measurements (at RHIC)

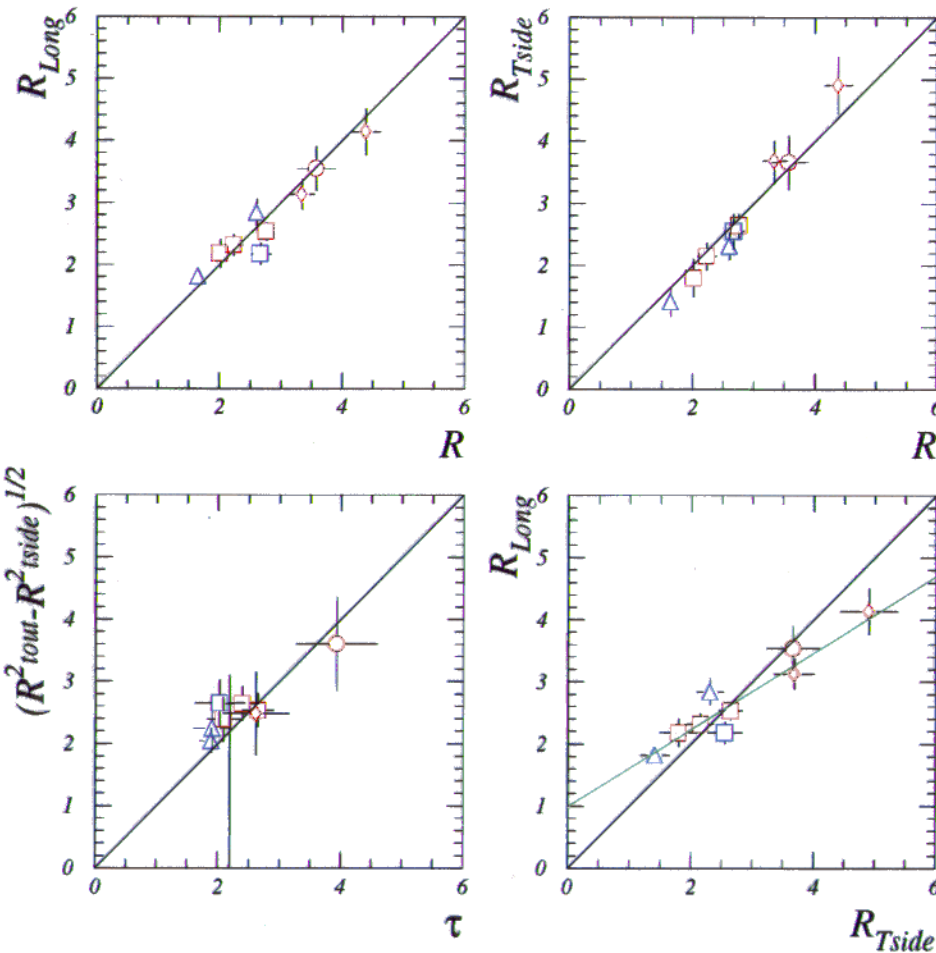
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Source Rapidity (AGS)



The 2D R - τ parameterization “works”!

E866 Soltz, Baker et al. Prelim.



R is a mixture of R_{Ts} & R_L

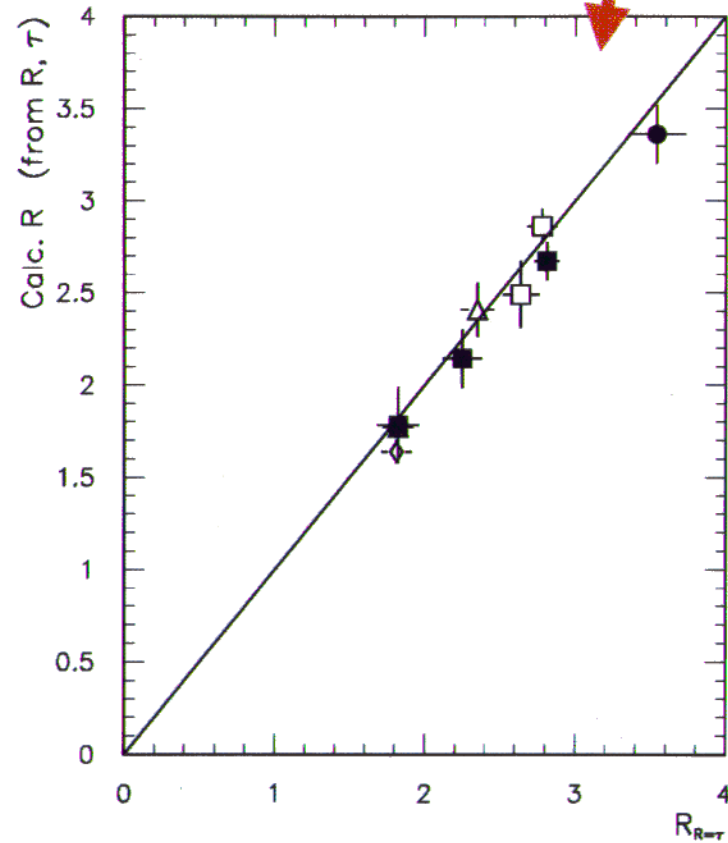
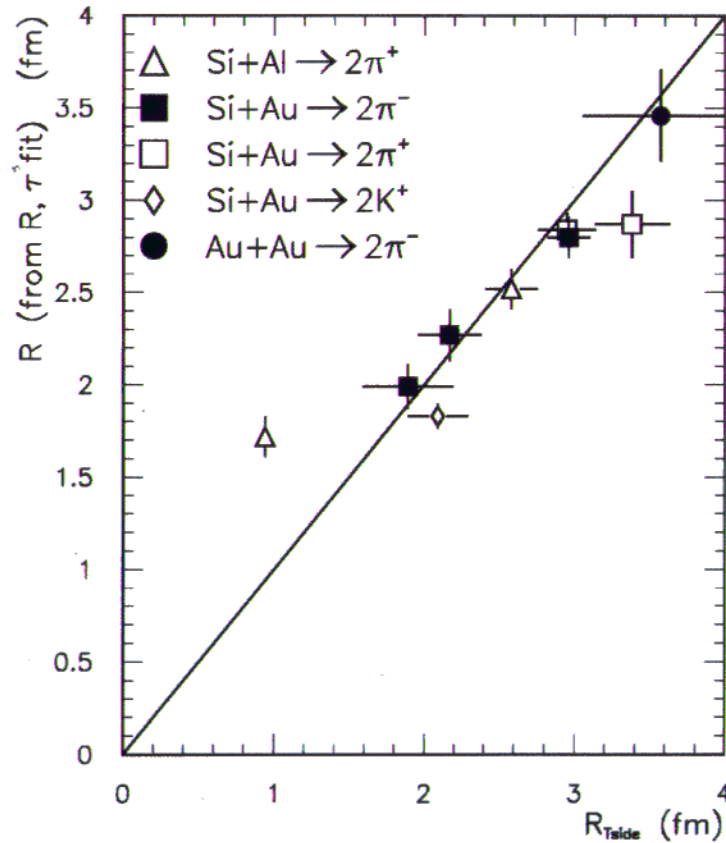
τ is equivalent to R_{diff}

The 1D R_{eff} parameterization “works”!!

$$Q_{\text{eff}}^2 = q^2 + q_0^2$$

$$R_{\text{eff}}^2 = 0.8R^2 + 0.2\tau^2$$

E859/E866 PRELIMINARY

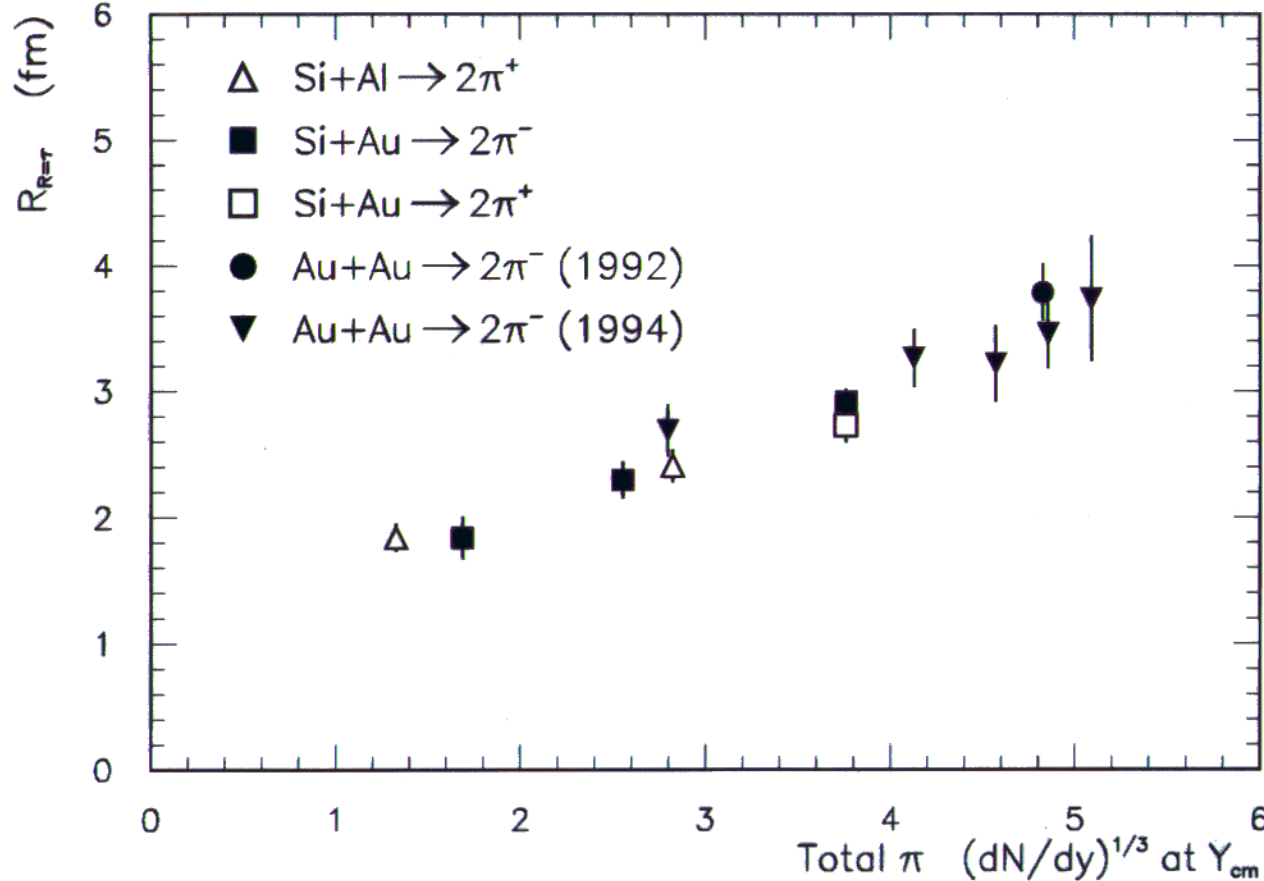


*E866
Baker, Soltz et al
Preliminary*

Centrality Dependence (1D)

No anomalous volume or duration increase

$$R_{1D} = \sqrt{0.8R^2 + 0.2\tau^2}$$



E866 Baker, Soltz et al.

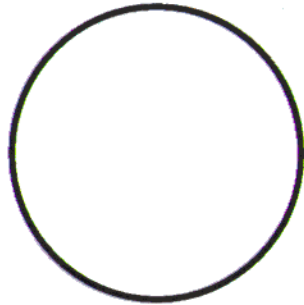
- Universal curve
- Good overlap.

Why HBT theory is non-trivial

A star (in space!)

\neq

A heavy ion collision



Static (time-independent)

Simple, well-understood geometry

Simple, well-defined surface

The source is "large" ($RP \gg \hbar$)

γ 's are independently produced (incoherent)

\vec{p} and \vec{x} are independent

γ 's arrive directly from the star

γ 's don't interact after emission

Dynamic (expanding, evolving)

Exact geometry unclear

Complicated "freezeout" surface

The source is "small" ($RP \gtrsim \hbar$)

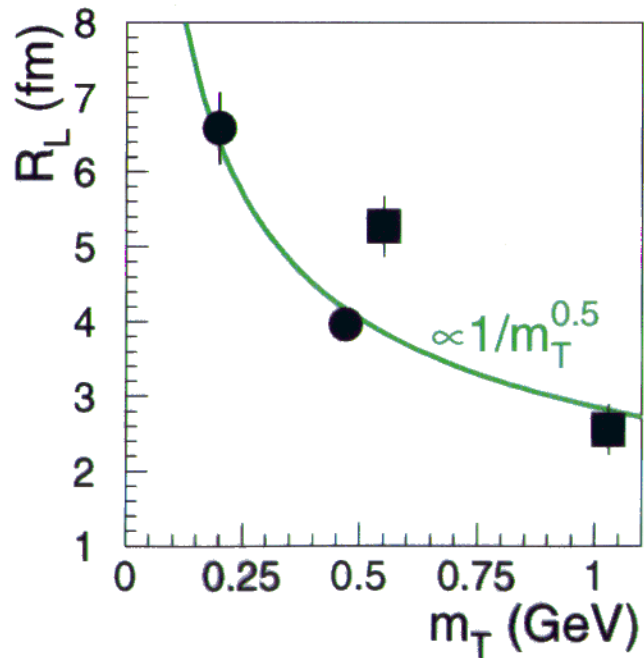
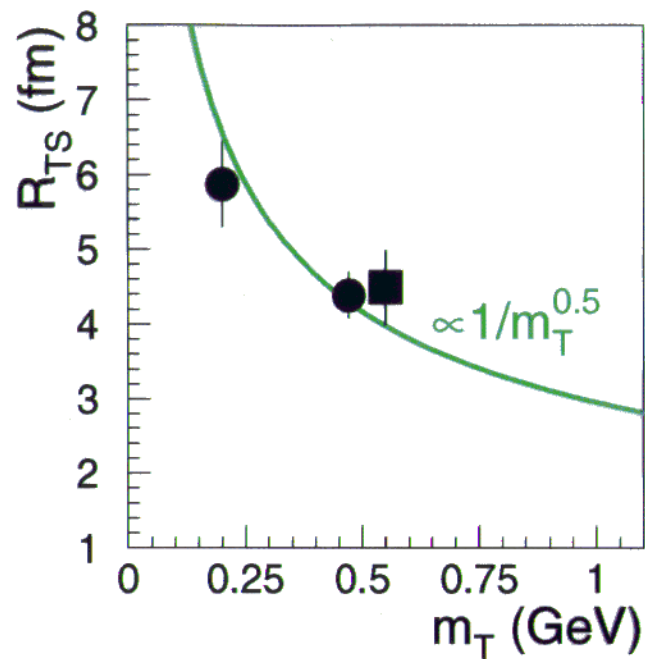
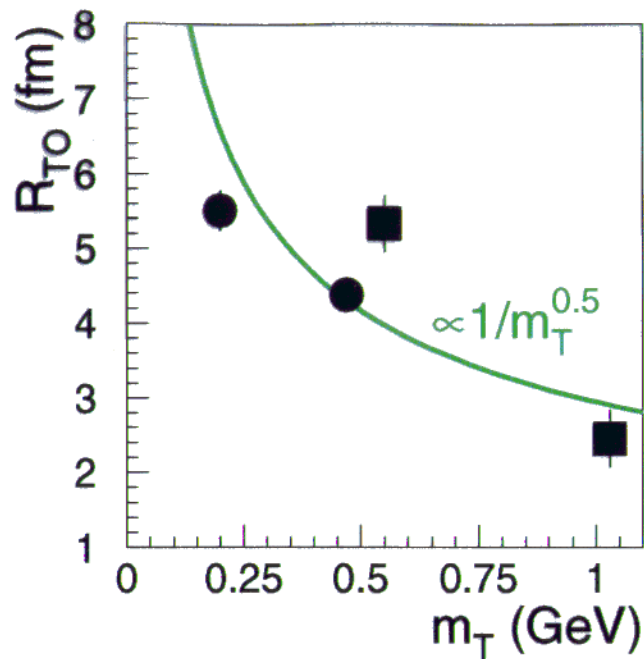
π 's may be coherent

\vec{p} and \vec{x} are correlated

Many π 's come from resonances

π 's interact after emission

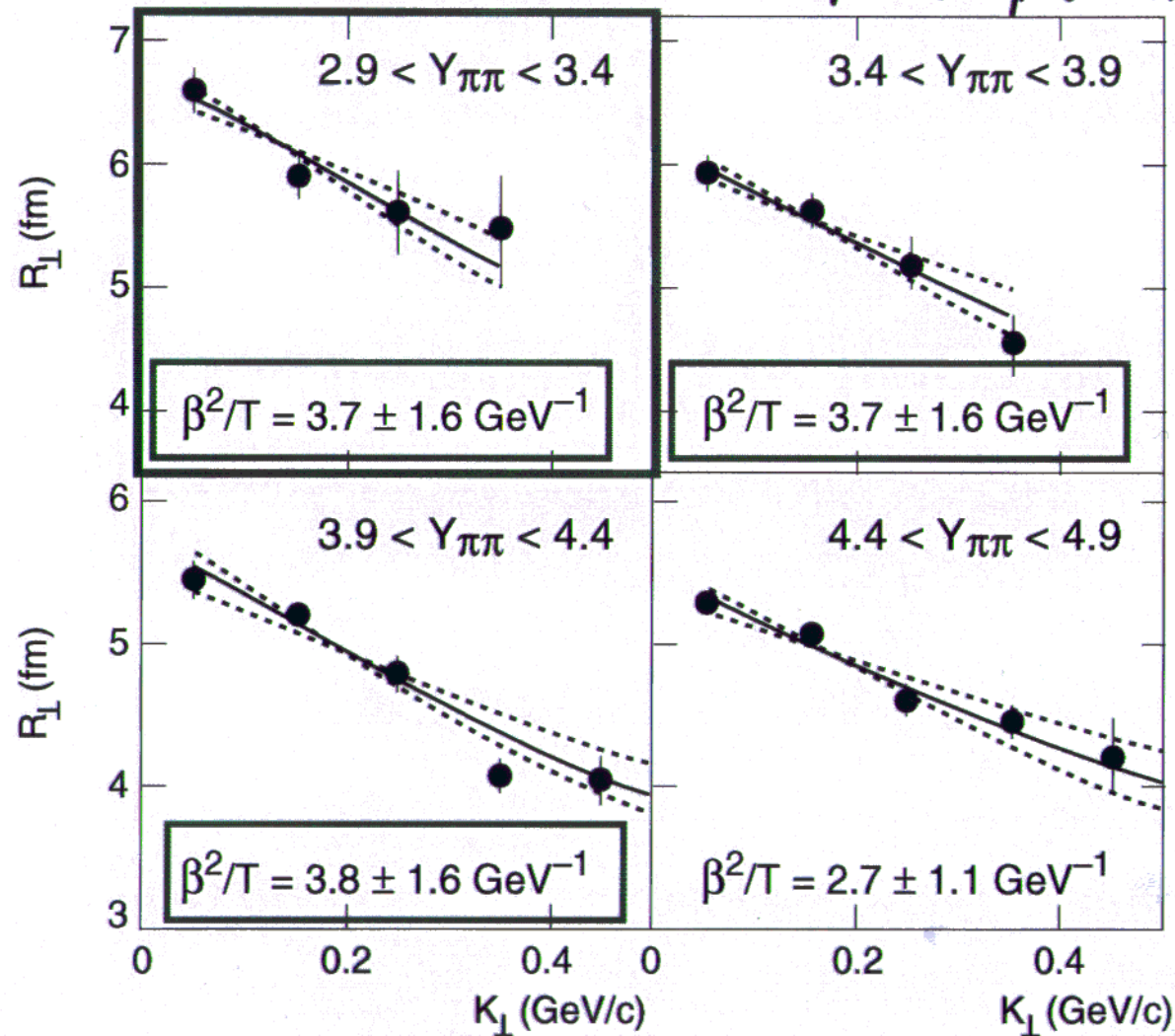
Pb+Pb at 158A GeV



- $Pb+Pb \rightarrow 2\pi^+ + X$
- $Pb+Pb \rightarrow 2K^+ + X$
(NA44 Preliminary)

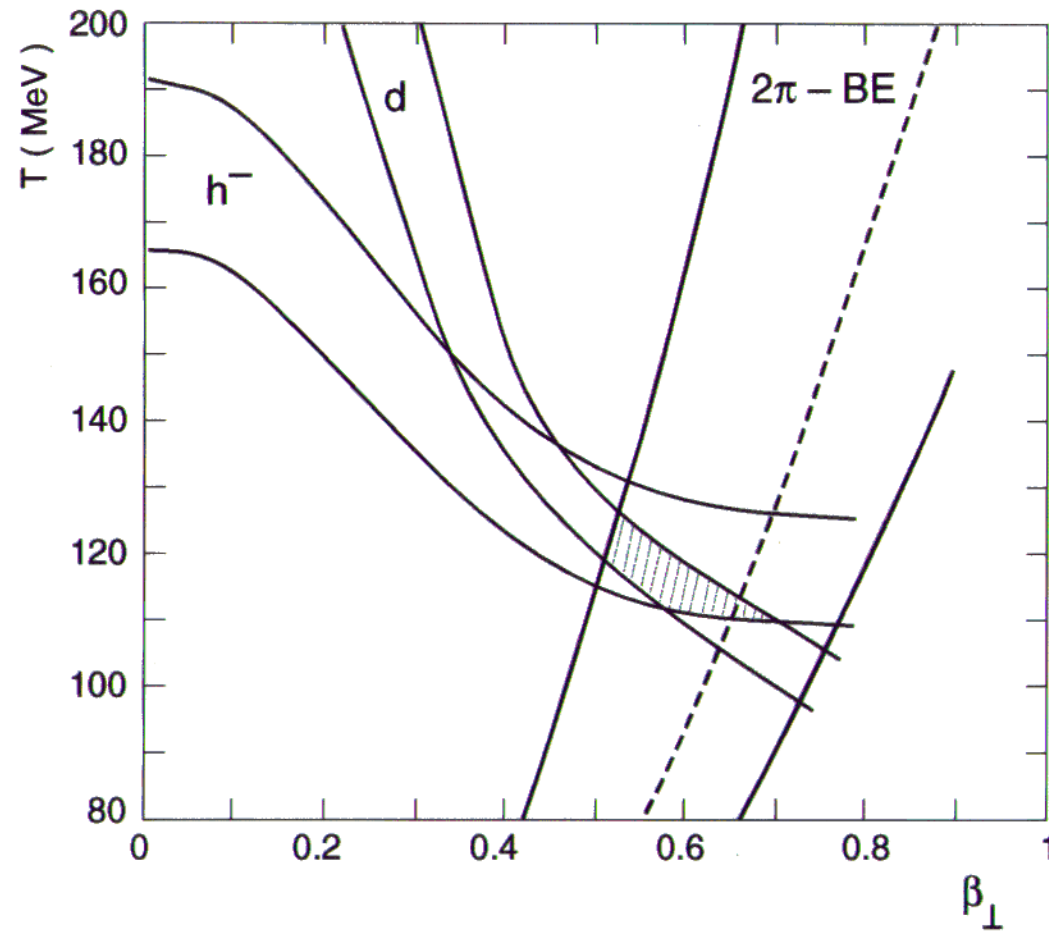
Transverse Expansion (SPS)

NA49 (prelim.?) PbPb



Combining HBT & singles info.

NA49 (Preliminary?)



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HBT Measurements (at RHIC)

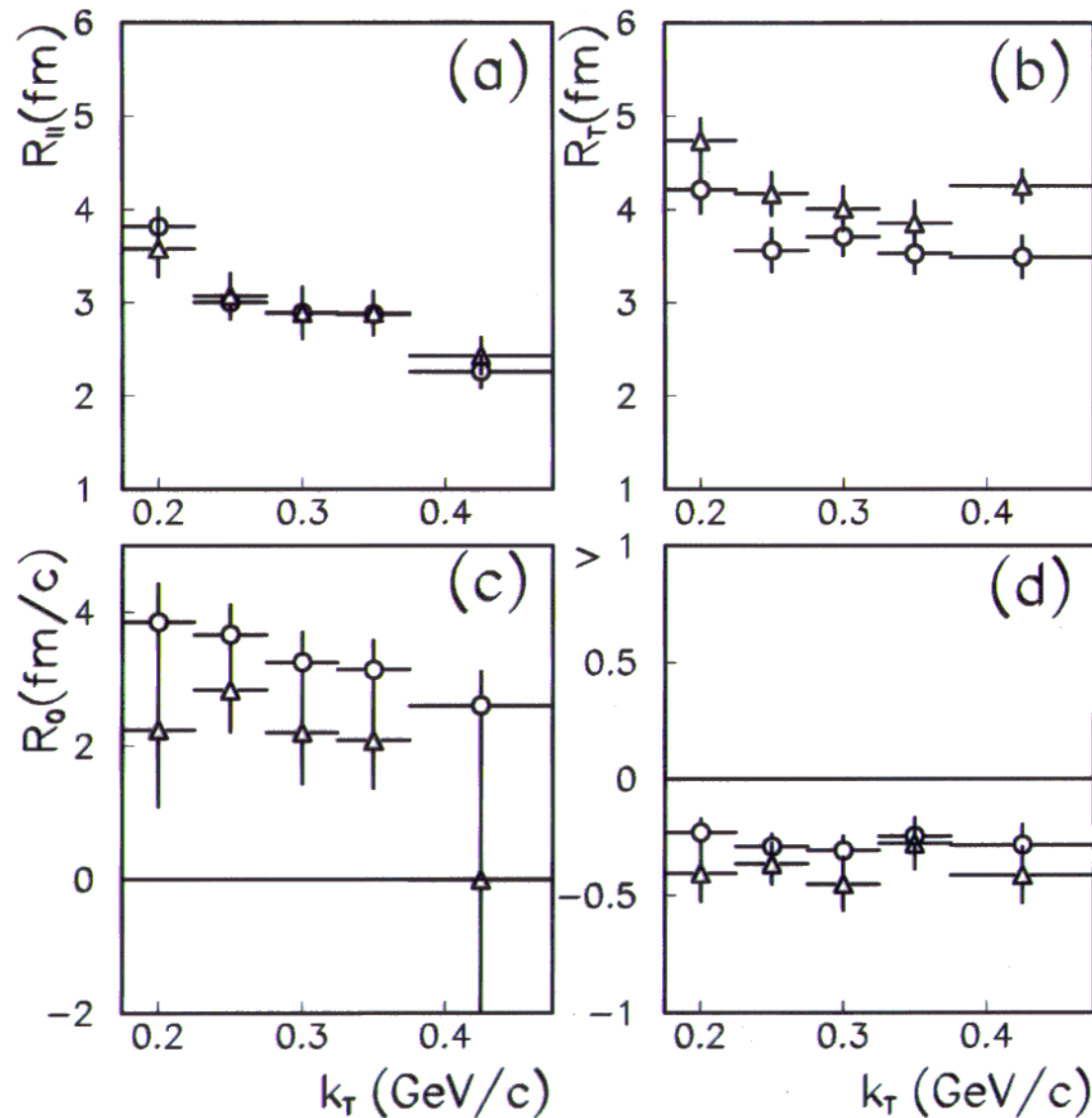
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AGS source: Near-mid-rapidity

E866 (Lee et al.)
Preliminary
AuAu

Δ $\pi^-\pi^-$
 \circ $\pi^+\pi^+$

$1.8 < y < 2.3$



JAN. 1999

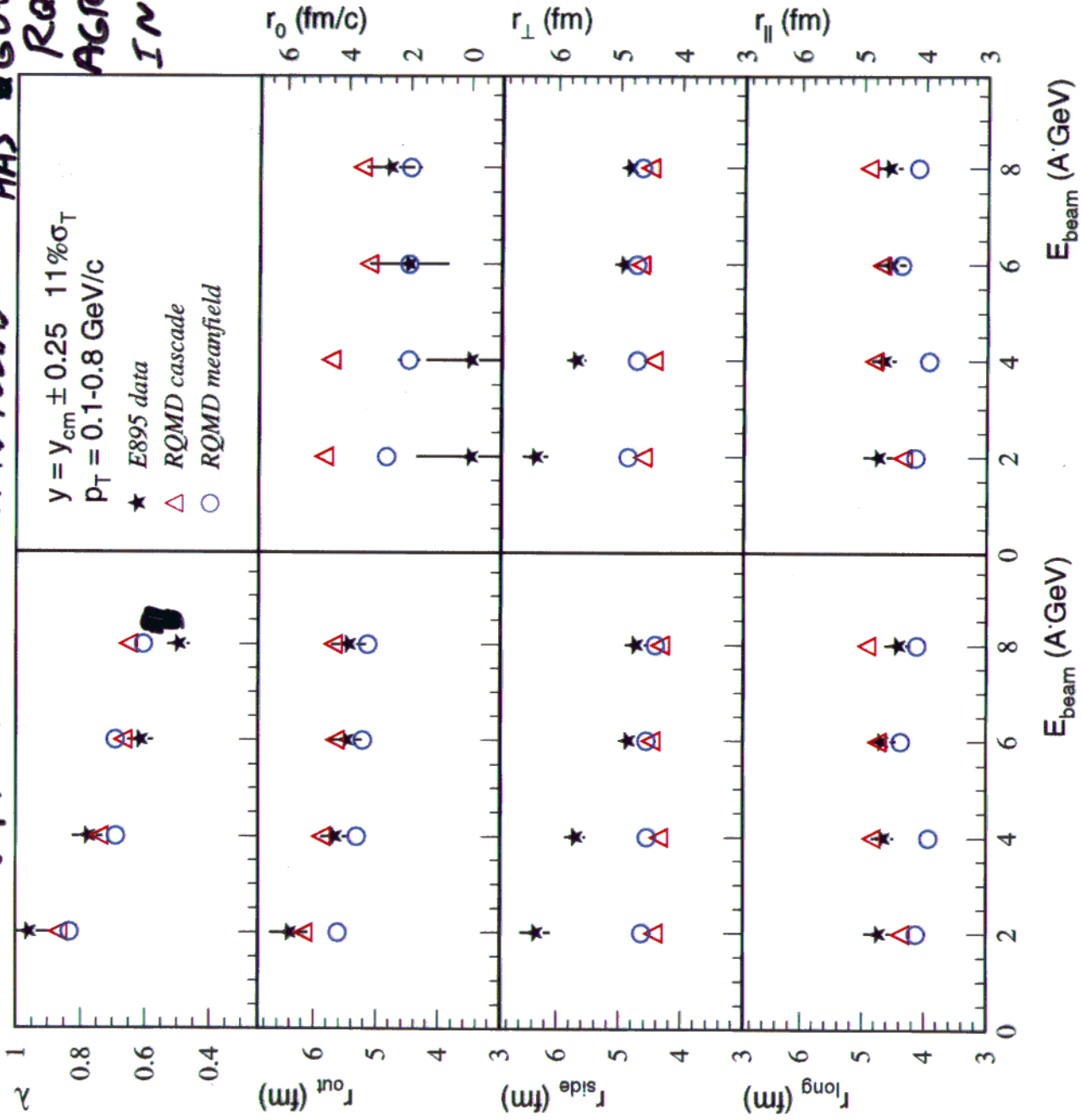
HBT Measurements (at RHIC)

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E895 (Lisa et al.) Preliminary

WARNING: OUT OF DATE SLIDE - CORRECT SLIDE HAS GOOD

RQMD AGREEMENT IN \nearrow

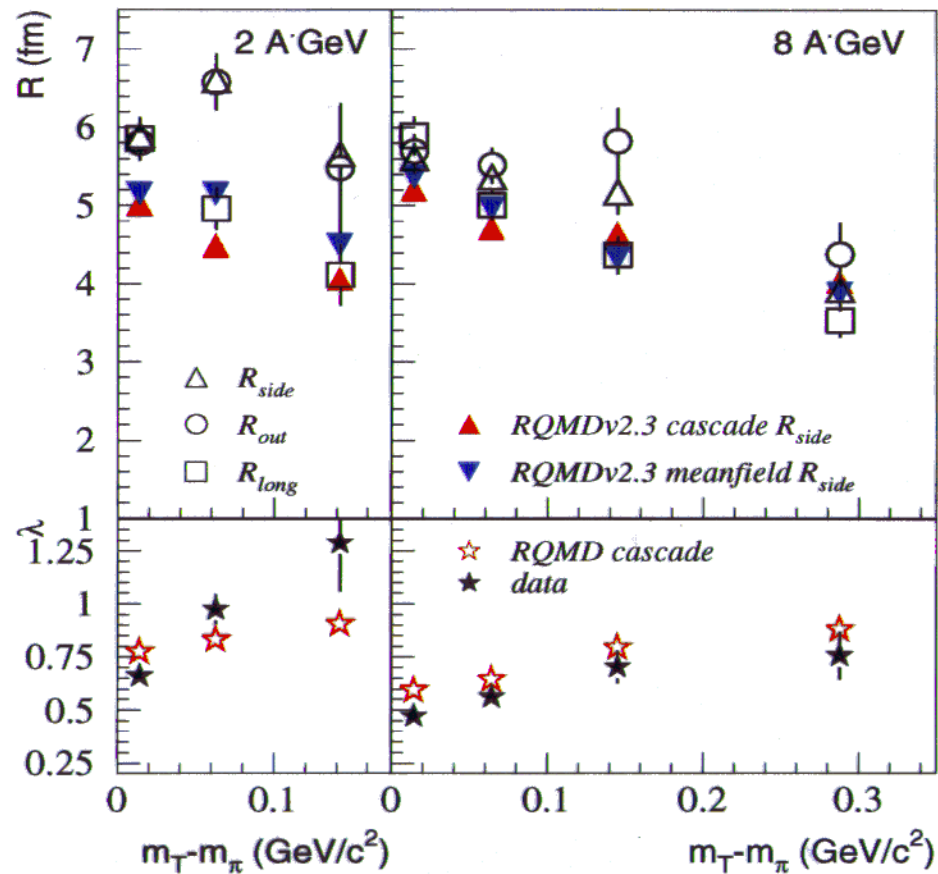


Why the large size (R_{T_S}) at low beam energy?

R_L and R_0 (duration of emission) are not large

E895 Kt dependence

E895 Lisa et al. Preliminary



PHENIX Simulations (CDR)

3.3. SELECTED EXAMPLES OF SIMULATIONS

3-49

STAR: Q_s even better

BRAMMS: $Q_s < 50 \text{ MeV}$

PHOBOS: " " for $l \text{ arm}$

Correlation Function of pions for $R=10 \text{ fm}$

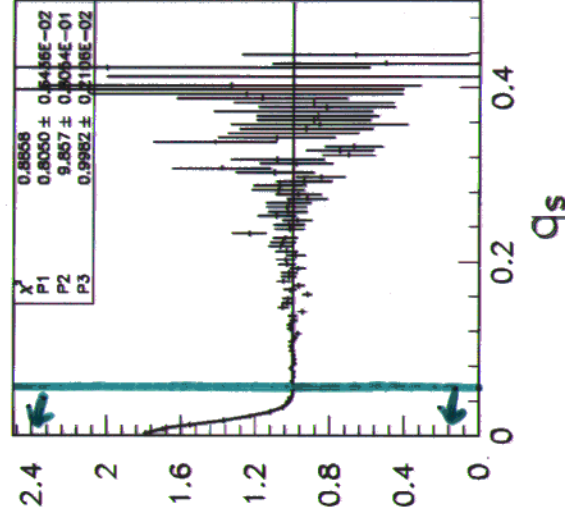
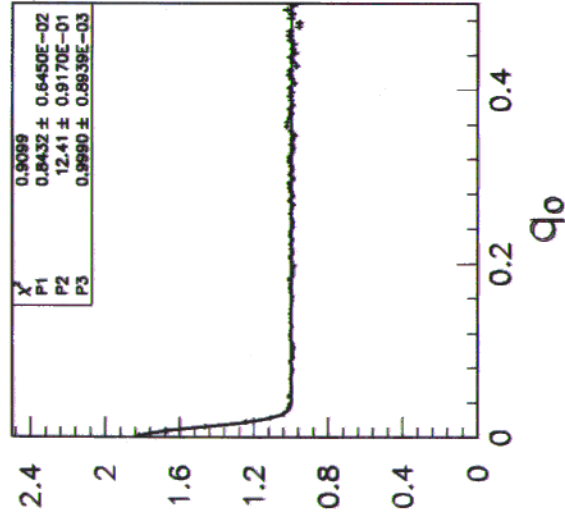
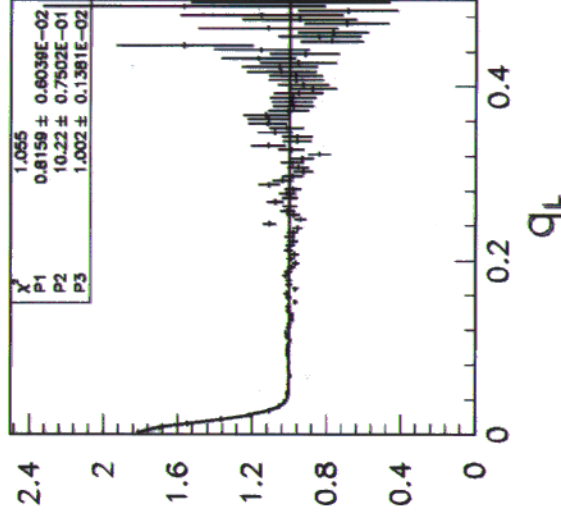
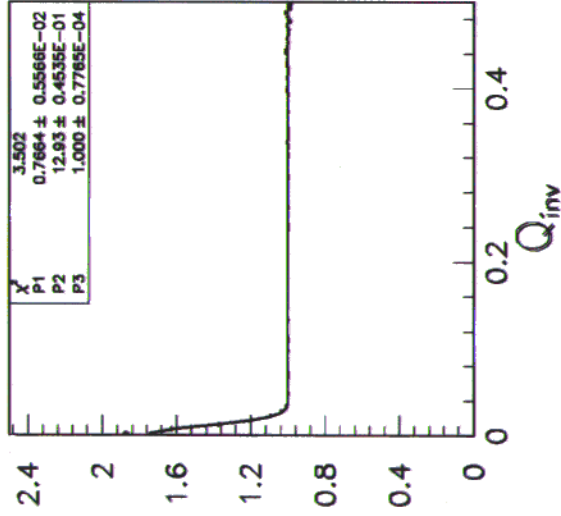
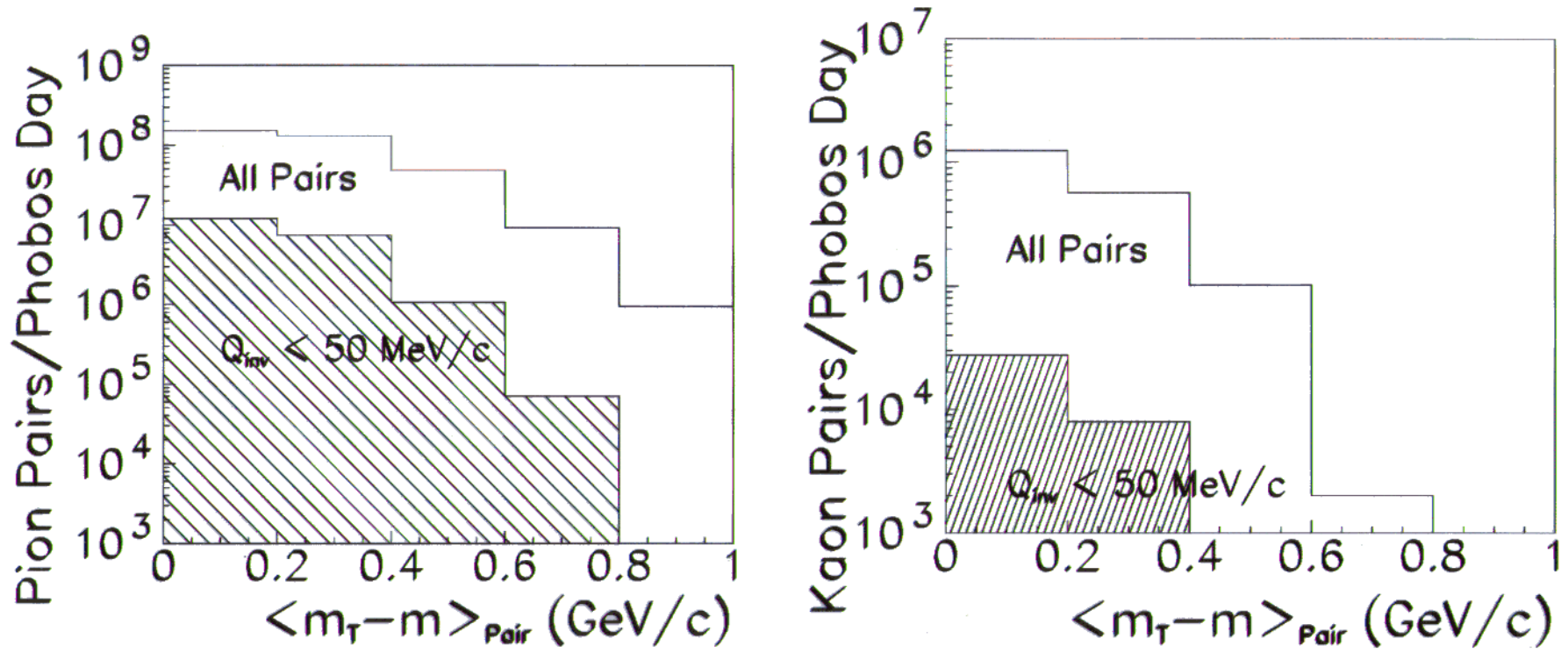


Figure 3.14: The correlation function of pions for a $R = \tau = 10 \text{ fm}$ source as seen in an aperture of $|\eta| < 0.7$, $\Delta\Phi = 30^\circ$ for 1 million Au + Au events.

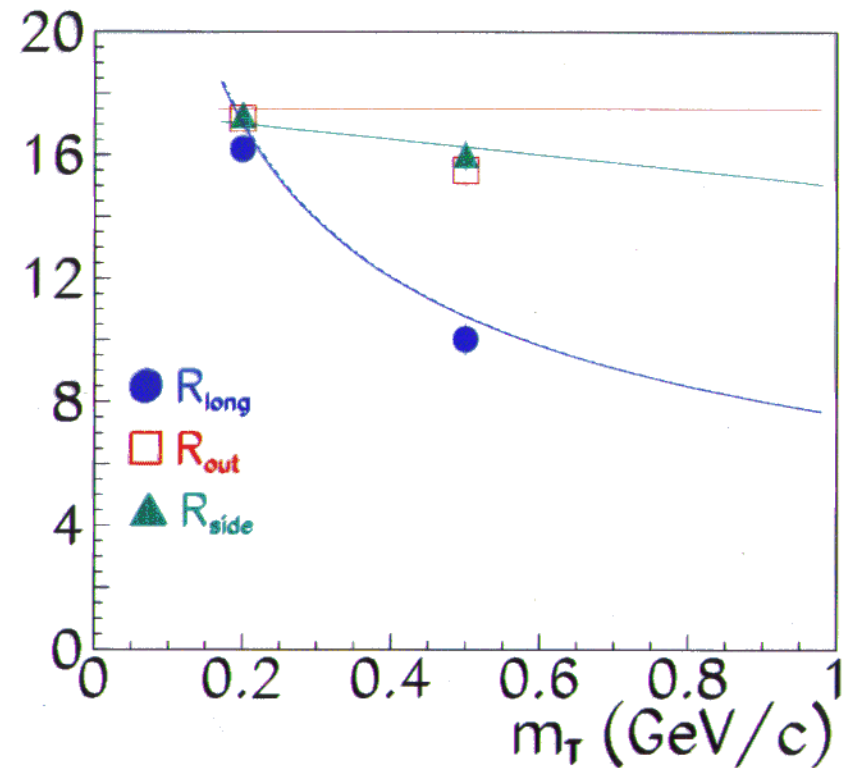
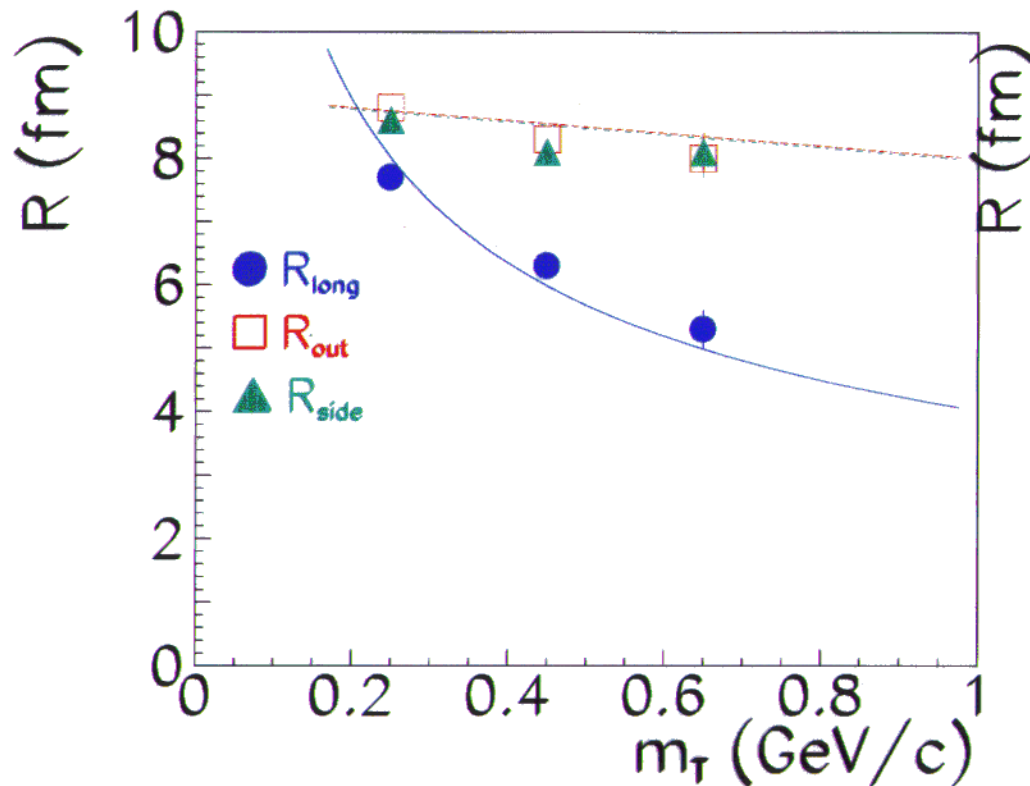
PHOBOS Rates (1 nominal day)



- 1 day nominal = the first 2-3 months in year 1
- Very high pair rates for all RHIC experiments
 - STAR gets ~ **a million per event!** (thousands at low Q)

PHOBOS HBT Capabilities

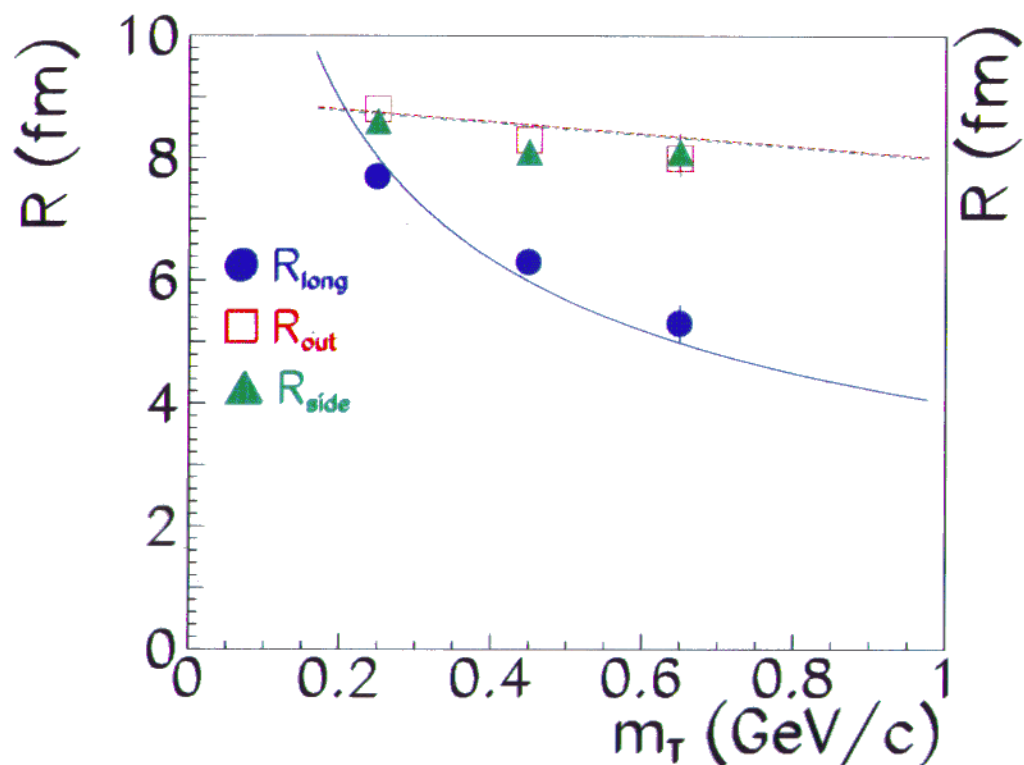
- “Standard” Scenario
 - Extrapolate from low E
- Exotic Scenario I
 - Very large source



PHOBOS HBT Capabilities

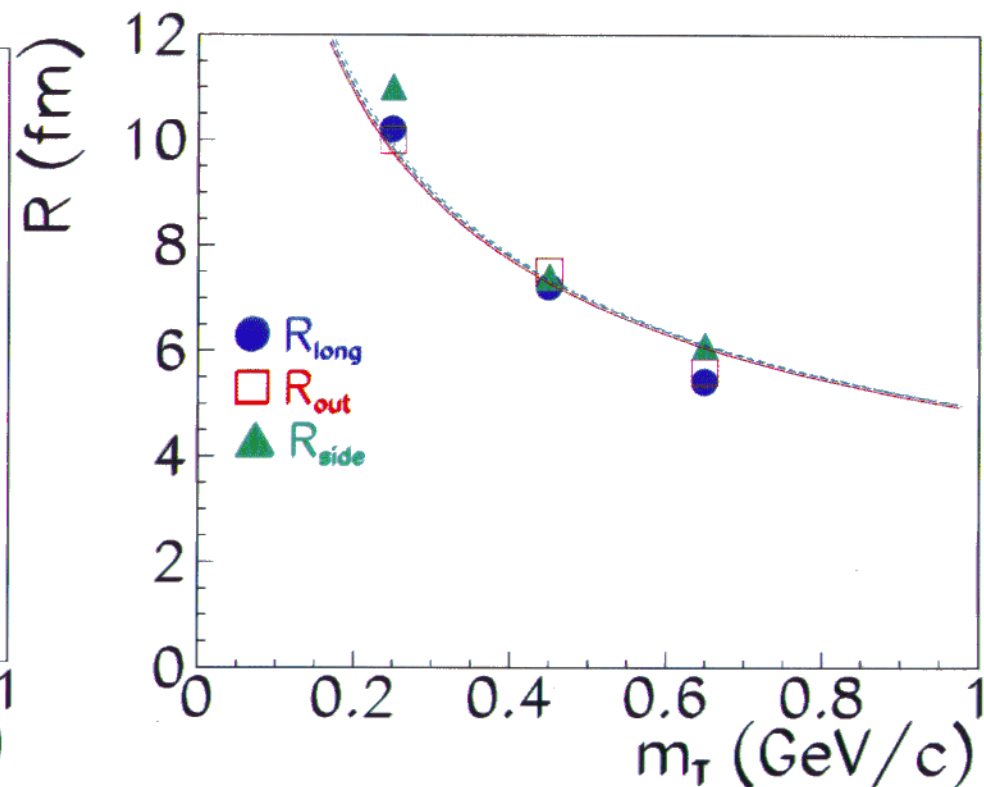
- “Standard” Scenario

- Extrapolate from low E



- Exotic Scenario II

- Very rapid expansion



Detectors @ RHIC

- BRAHMS:
 - Pro: Very broad (patchwork) coverage to high y , p_T
 - Con: Acceptance problem for small sources
- PHENIX:
 - Pro: Good acceptance
 - Con: Low p_T particles decay before being PID'd
- PHOBOS:
 - Pro: Good acceptance at low p_T
 - Con: Need 2 arms for small sources
- STAR:
 - Pro: Very large acceptance - full azimuth
 - Con: Annoying high Q combinatoric background
 - NOTE: Event-by-event will be crude

Conclusions

- HBT has matured
 - The theory has improved dramatically!
 - We've learned a lot at the AGS & SPS
 - We stand to learn more at RHIC
- More to do (theory)
 - Understand 1D & 2D parameterizations
 - there will ALWAYS be low statistics data